# Dwarf galaxy mergers





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t = 3.34

### Why merging dwarf galaxies?

- ACDM cosmology: Hierarchical structure formation
- Simulations of isolated galaxies:(Valcke et al. 2008, Schroyen et al. 2010)
  - $\rightarrow$  making simulations more cosmologically motivated by adding a merger history
- Merger trees:

Millenium run (Springel et al., 2005)



### Simulations

#### Code: modified version of Gadget2 (Springel et al. 2005)

- + star formation
- + feedback
- + cooling (metallicity dependent radiative cooling (Sutherland and Dopita, 193) and cooling below  $10^{4}$ K (Maio, 2007))

#### **Initial conditions:**

- $\rightarrow$  Joining 2 isolated simulations
- $\rightarrow$  Orbital parameters of galaxies  $\rightarrow$  Benson et al., 2005

### **Isolated simulations**

#### **Initial setup:**

- spherical symmetric dark matter (DM) halo with NFW density profile
- homogeneous gas cloud



 stable in DM only simulations
 conversion from cusp to core in DM + gas simulations and in DM + gas + star formation simulations.

(Navarro, Frank & White, 1994)

#### Star formation criteria:

$$egin{array}{lll} ec 
abla .ec v &\leqslant 0, \ 
ho_{
m g} &\geqslant 
ho_{
m SF} \end{array}$$

### **Isolated simulations**

**Feedback efficiency**  $\boldsymbol{\epsilon}_{FB}$  = fraction of supernova energy that is absorbed by the ISM (SNIa, SNII and stellar winds)

**Other poster:** parameter survey to investigate the degeneracy

between  $\boldsymbol{\epsilon}_{FB}$  and  $\boldsymbol{\rho}_{SF.}$ 

$$\begin{array}{c} \rho_{\rm\scriptscriptstyle SF}=6\ {\rm cm}^{{\rm\scriptscriptstyle -3}}\ \rightarrow \epsilon_{\rm\scriptscriptstyle FB}=0.5\\ \rho_{\rm\scriptscriptstyle SF}=50\ {\rm cm}^{{\rm\scriptscriptstyle -3}}\ \rightarrow \epsilon_{\rm\scriptscriptstyle FB}=0.7 \end{array} \end{array}$$

Simulations are in agreement with the observed kinematical and photometric scaling relations



**Top**: half-light radius versus V-band magnitude **Right**: metallicity versus V-band magnitude







20

40

60

t = 2.80

10

1e-01

1e-02

1e-03

densitu

20

-60

-80

-100

-120

-20

100 000 gasparticles 100 000 DM particles  $\rightarrow$  total mass: 1.5 10<sup>9</sup> M<sub>sol</sub>



Increase of the star formation rate at the moment of the



Final galaxy is **triaxial** and is **rotating** 

#### **Small gas clumps:**

We observe very small gas clumps which collapse and in which stars form. Supernova explosions redistribute the gas and create a bubble structure in the ISM.



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 $\rightarrow$  kinematic and photometric scaling relations  $\rightarrow$  slope of the  $M_{_{star}}\text{-}M_{_{halo}}$  relation



Relation between the stellar and halo mass

## Movie

#### Thank you for your attention. Questions?