INTERSTELLAR MOLECULES: DIRECT FORMATION ON GRAPHITE GRAINS AT 7-78K

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Abstract. The formation of CH_4 and C_2H_6 by the reaction of H-atoms with graphite when both are at $T>7\,\mathrm{K}$ was demonstrated experimentally. In another set of experiments at $78\,\mathrm{K}$, H-atoms produce CH_4 and C_2H_6 , O-atoms produce CO and CO_2 and S-atoms produce CS_2 . No CN species were formed by N-atoms. Implications to interstellar clouds are discussed.

1 Introduction

The theoretical aspects of the recombination of H-atoms on cold grain surfaces has been studied extensively over the years. Apparently, the problem has not been solved completely, as can be seen in recent theoretical studies by Katz et al. (1999) and by Cazaux and Tielens (2004). Three recent experimental studies on recombination of H-atoms to form H_2 molecules were carried out by Pirronello et al. (1999) on amorphous carbon, by Hornekaer et al. (2005) on graphite and amorphous water ice and by Perets et al. (2005) on amorphous water ice. The chemistry on interstellar grains has been reviewed recently by Herbst et al. (2005). Yet, in all the above mentioned studies and in all others, it has not occurred to the authors that H atoms, in addition to recombining on graphite or amorphous carbon, can react with it to form hydrocarbons. Obviously, all the surfaces of graphite

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and amorphous carbon are unsaturated and are prone to reactions with incoming atoms. The reactions of H-atoms with graphite down to 7 K were shown already in 1980 to produce CH₄ and C₂H₆ (Bar-Nun et al. 1980). At 78 K CH₄ and C₂H₆ were produced, as well as CO and CO₂ with O-atoms and CS₂ with S-atoms (Bar-Nun 1975). Although no new experimental results became available in the meantime, I would like to draw attention to this mechanism of molecules formation in interstellar clouds and its consequences.

2 Experiments with H atoms at 4.6-300K

A detailed description of the experiments is given in Bar-Nun et al. (1980). Therefore, only a brief description is given here. The 50 cm long, 2 cm inside diameter Spectrosil quartz reaction vessel is shown schematically in Fig. 1. For final removal of all remaining traces of organic contaminants, the vessel was filled with pure oxygen and heated to 1200 K, so that the organics would be oxidized to CO_2 and be removed by pumping to 8×10^{-6} Torr at 1200 K. This burning procedure was performed twice.

In the blank runs, without graphite, either hydrogen (Matheson, Research Purity, 99.9999% pure) or helium (Matheson, Research Purity, 99.9999% pure) were introduced at pressures of 6 Torr into the reaction vessel. A microwave discharge, generated by a 100 W Kiva model MPG-4 microwave power generator, was initiated in the 10 cm long side arm of the reaction vessel. The discharge was visible only in the side arm and did not appear to extend beyond it into the reaction vessel. After 60 min. the discharge was stopped, 200 Torr of either hydrogen or helium was introduced into the vessel and allowed to mix thoroughly with the products for 24 hrs. Analyses of the gas mixture were done on a Packard-Becker model 417 gas chromatograph, using a flame-ionization detector. Pieces of pyrolytic graphite (General Electric, 3 mm long and 0.5 mm thick) were introduced to the bottom of the vessel and pumped to 8×10^{-6} Torr at 1200 K for 2-3 hours. After the vessel was cooled to room temperature, 6 Torr of helium or hydrogen and helium at a pressure of 6 Torr each were introduced. The vessel was then inserted into the neck of a liquid helium dewar. The graphite temperature was monitored by a carbon resistor, (precalibrated by comparison with a germanium cryogenic thermometer) which was inserted into an inner quartz tube filled with helium. After equilibration at the desired temperature, the microwave discharge was switched on in the side arm and

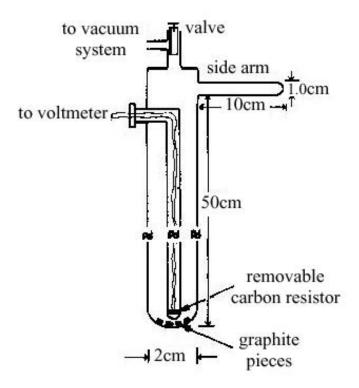


Fig. 1. A schematic drawing of the reaction vessel (not drawn to scale)

continued for 60 min. At the onset of the discharge, an increase in temperature of between $0.2\text{-}0.4\,\mathrm{K}$ was observed. At the end of the experiment, the vessel remained in the dewar for at least 5 min. and then warmed to room temperature, 200 Torr of helium were introduced and the analysis was done 24 hrs. later, to ensure thorough mixing. The experimental conditions and product distribution are presented in Table I.

From Table I, it is clear that $\mathrm{CH_4}$, the major product, is still produced by H-atoms impinging on graphite, when both are cooled to 7 K. At 4.6 K, the yield of methane is similar to that of the blank run with hydrogen in the absence of graphite, which suggests that at 4.6 K hydrocarbons are not

Table 1. Product Distribution in the Various Experiments

Run		Reacta	nt	Temperature (K)	Products- nmole after 60 min. of discharge				
	H_2	He	graphite		CH ₄	C ₂ H ₆	C ₂ H ₄	C ₂ H ₂	
21	+			300	2.35				
22	+	+		300	2.38	0.21	0.21		
23		+	+	300	0.19				
30		+	+	300	0.93			0.08	
32		+	+	300	0.33				
45		+	+	300	0.30		0.08		
35	+	+	+	300	10.37				
46	+	+	+	78	29.46	0.59	0.22		
25	+	+	+	20	45.77	0.59	0.09	0.23	
37	+	+	+	12	62.20	0.75	0.58		
38	+	+	+	12	65.65	0.52	0.32		
44	+	+	+	9	7.62	0.26			
47	+	+	+	9	7.44	0.07	0.52		
41	+	+	+	7	5.20				
42	+	+	+	7	4.60	0.05	0.48		
40	+	+	+	4.6	1.60	trace	0.27		

formed. As the temperature is raised, the yield of hydrocarbons increases up to 12 K and then decreases. Two effects are probably responsible for this behavior: at very low temperatures, hydrogen is adsorbed on the graphite until, finally, at 4.6 K no hydrogen is left in the discharge zone. On the other hand, higher hydrogen pressure results in a smaller mean free path for the Hatoms and their subsequent gas phase recombination to form H₂ molecules. The interplay between these two effects may be responsible for the peak in production at 12 K. Because of the adsorption of hydrogen on the graphite, 6 Torr of helium were introduced into the vessel along with the hydrogen, to serve as a heat bath. Although helium is also adsorbed on graphite, it does so to a lesser extent and even at 4.6 K its pressure was higher than 0.5 Torr. This is more than enough to ensure thermalization of the H-atoms through collisions, during their diffusion along the 50 cm path from the discharge to the graphite. Table I also shows that C_2H_6 at about a 10th of the CH_4 is formed down to 12 K, with some traces of C₂H₄ and perhaps C₂H₂ at 20 K. A question might arise as to whether the hydrocarbons could have been formed by trapped H-atoms in the graphite only when the sample was warmed up to room temperature at the end of the experiment. This is very unlikely, since in our experiments more than 5 min. passed (intentionally) from the time the RF-discharge was switched off and the vessel was removed from the liquid helium dewar. By this time, all the H-atoms trapped in the graphite should have recombined to $\rm H_2$ and desorbed from the graphite since, as found by Pirronello et al. (1999), hydrogen molecules desorb from the amorphous carbon within tens of seconds at 15-17 K. The ratio of the probabilities for H-atom recombination on a graphite surface vs that for H-atom reaction by quantum-mechanical tunneling with the graphite to yield $\rm CH_4$ should be:

$$P_{recomb}/P_{react} = \pi a^2 < v > sn_H/\nu_0 exp[-(4\pi/h)(2m\Delta E)^{\frac{1}{2}}\Delta x]$$

where the recombination was calculated by Barlow and Silk (1976), after Hollenbach (1969) and Hollenbach and Salpeter (1970). a is the grain radius, $\langle v \rangle$ is the mean thermal velocity of the gas-phase H-atoms, s is the H atom sticking probability (\sim 1). $n_{\rm H}$ is the gas-phase H-atom density and m is its mass. $\nu_0=10^{13}\,sec^{-1}$ is the frequency of surface vibration of an adsorbed H-atom and $\Delta E=5500$ cal mole⁻¹ is the activation energy of the reaction of an H-atom with graphite (Wood & Wise 1969). This yields at \sim 30 K $P_{recomb}/P_{react}=1.3\times10^{-10}\,\rm n_{H}$. Under our experimental conditions, a roughly estimated upper limit of $\rm n_{H}\gg10^{12}~cm^{-3}$ near the graphite would result in an estimated $P_{recomb}/P_{react}\gg130$, which would still allow a measurable rate of methane formation. With $\rm n_{H}=10^{3}~cm^{-3}$ in clouds, $P_{recomb}/P_{react}\gg1.3\times10^{-7}$. Namely, the major pathway will be hydrocarbon formation rather than recombination. This will be discussed later.

3 Experiments with H, O and S-atoms at 78 K

A detailed description of the experiments is given in Bar-Nun (1975). Graphite powder 170-300 mesh, "specpure" from Johnson Matthey was used and the vassel was cooled by liquid nitrogen. H and O-atoms were produced by an RF discharge (Raytheon PGM-10) in H₂ and O₂ as described before. S-atoms were produced by a discharge in sulfur vapor from sublimation of sulfur powder ("sublimed" from Riedel-de Haen) at the end of the side arm, in the presence of 3 Torr of He. Some flow experiments were carried out with H and O-atoms, but will not be discussed here. The experimental results are shown in Table II.

Table 2. Product Distribution in the Various Experiments (static RF series)

Run	React. gas	P (Torr)	Temp (K)	Time (min)	Products- µmole							
					CH ₄	C ₂ H ₆	C ₃ H ₈	C ₄ H ₆	C ₄ H ₁₀	CO	CO ₂	CS ₂
1	Ar	3	300	150	2.4-3							
2	H_2	3	78	60	6.8 - 1	2.0-2	<1.0-4	7.7-4	8.3-4			
3	H_2	30	78	60	2.1-3	2.3-4						
4	H_2	3	300	60	1.1-1	1.3-3	1.5-3	1.1-3	1.7-2			
5	O_2	3	78	180						4.5-1	3.7	
6	S	3(He)	78	90								4.5

Again, the major hydrocarbon product is CH_4 with about a tenth of C_2H_6 and some C_3H_8 and C_4H_{10} . As discussed by Bar-Nun (1975), the C_3 and C_4 species are produced most likely in the gas phase at $78\,\mathrm{K}$, while CH_4 and C_2H_6 are produced on the graphite surface. Similarly, CO, CO_2 and CS_2 are produced on the graphite surface. N-atoms did not produce C_2N_2 , which could have been expected by analogy to C_2H_6 .

4 Discussion

The major conclusion of these experimental studies is that in dense clouds where $n_{\rm H}=10$ –1000 cm⁻³, $P_{recomb}/P_{react}=1.3\times10^{-9}$ –1.3 \times 10⁻⁷ and the H-atoms will react preferably with bare graphite surfaces to produce mostly CH₄, rather than recombine to form H₂. The CH₄ and some C₂H₆ would stay on T < 20 K surfaces and form a monolayer, on which further H-atom recombination to H₂ could proceed. Over time, UV photolysis and low energy cosmic rays could polymerize the hydrocarbons coating, leading to high molecular weight materials similar to Titan's Tholins (Bar-Nun et al. 1988), as well as eject some of the small molecules. In clouds at T > 78 K, CH₄, C₂H₆, CO, CO₂ and, perhaps, CS₂ would desorb from the grain surface after their formation, leaving behind bare carbon surfaces. The Pirronello-Vidali group (1999) study of H-atom recombination on carbonaceous material is an important step in the right direction. I would like to encourage them to study as well the formation of CH₄ and other hydrocarbons under the same conditions.

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