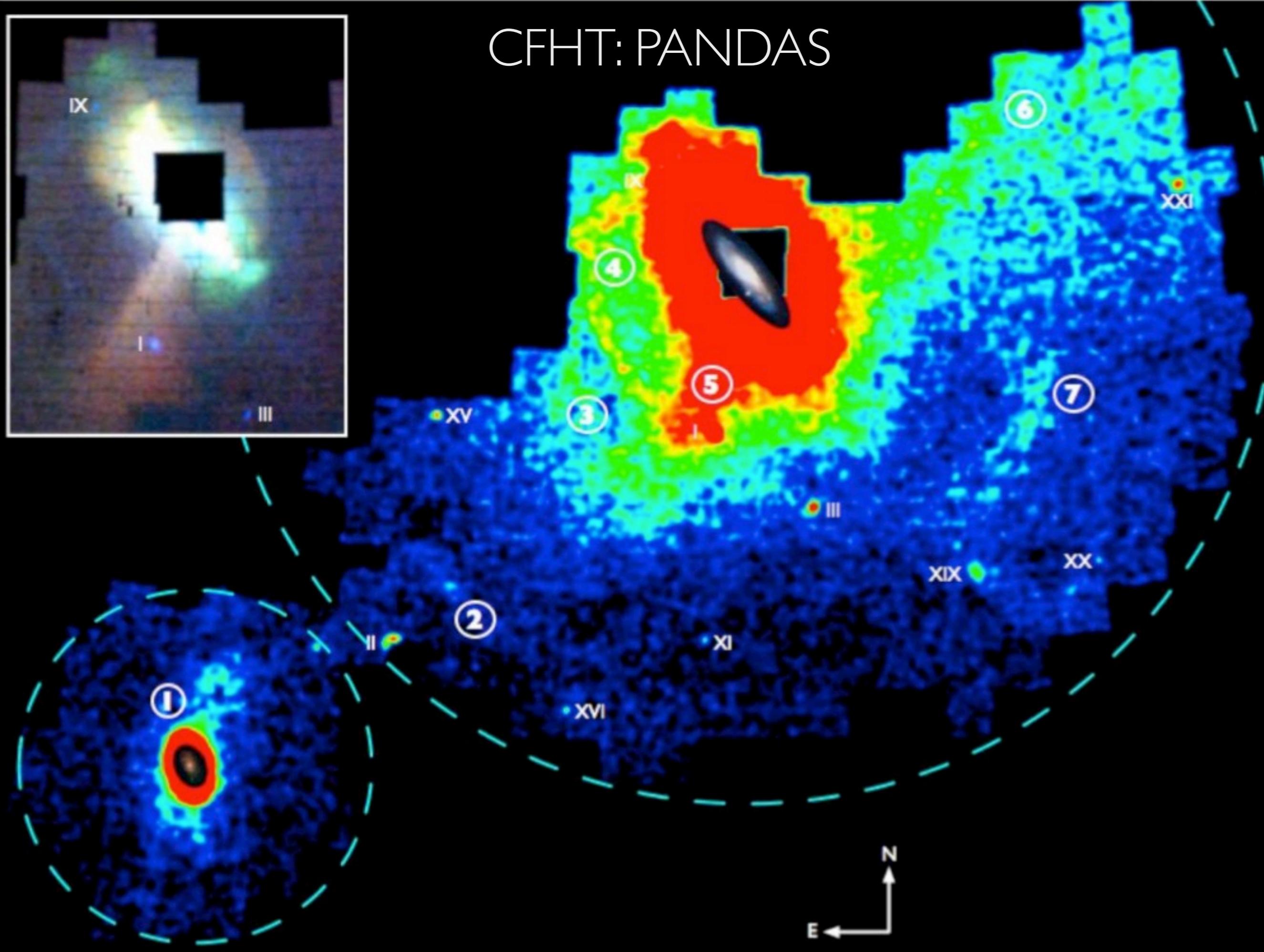
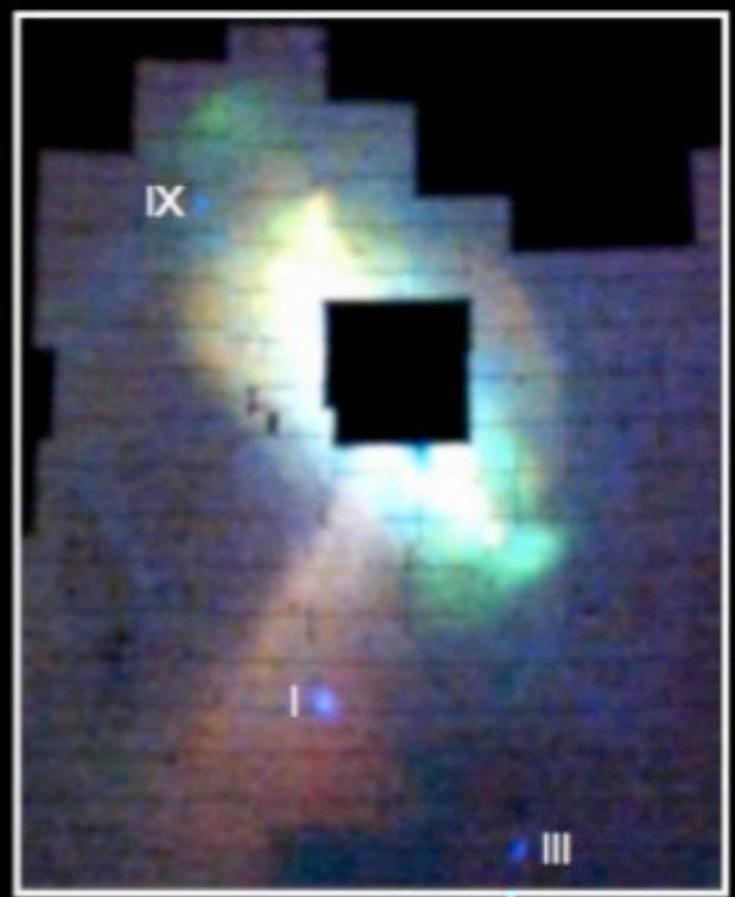


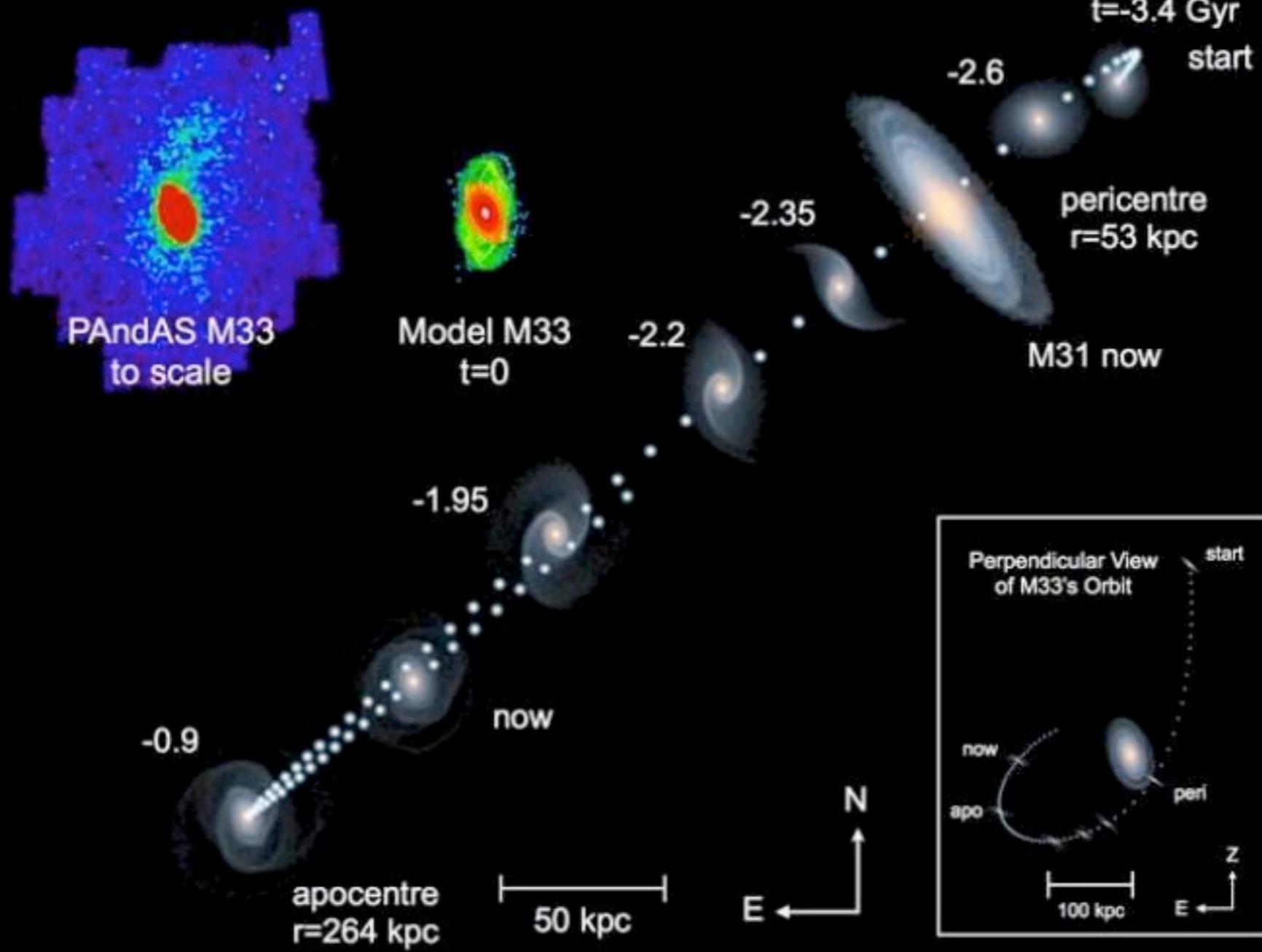
SKA PRECURSORS

Galactic accretions...

CFHT: PANDAS

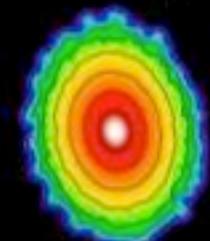


An M31-M33 Interaction Model



Surface Brightness

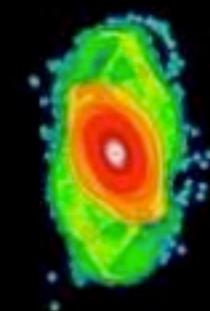
$t=-3.4$ Gyr



μK

35

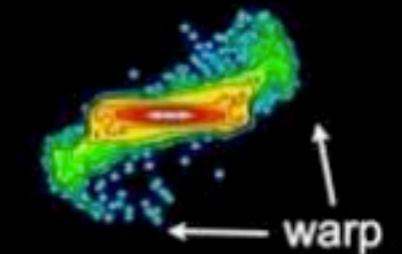
$t=0$



30

25

Edge-on $t=0$



20

18

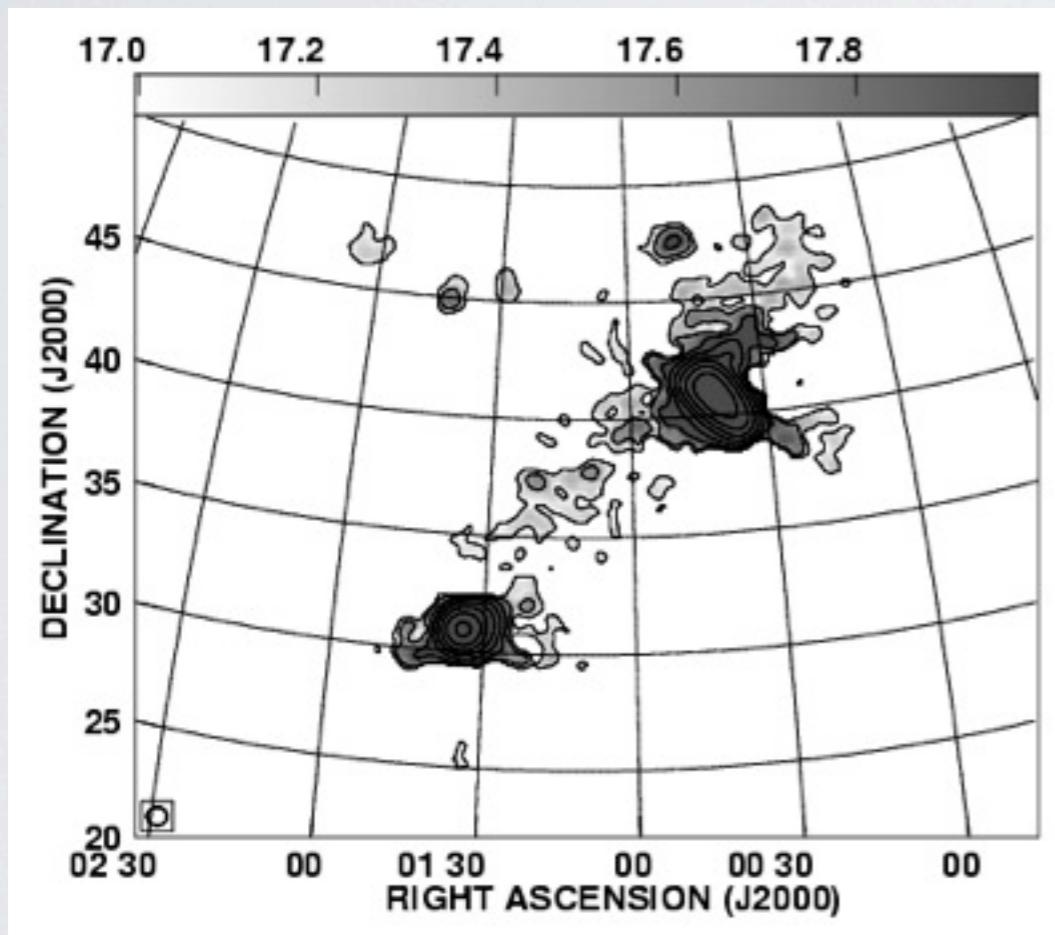
20 kpc

McConnachie et al 2009

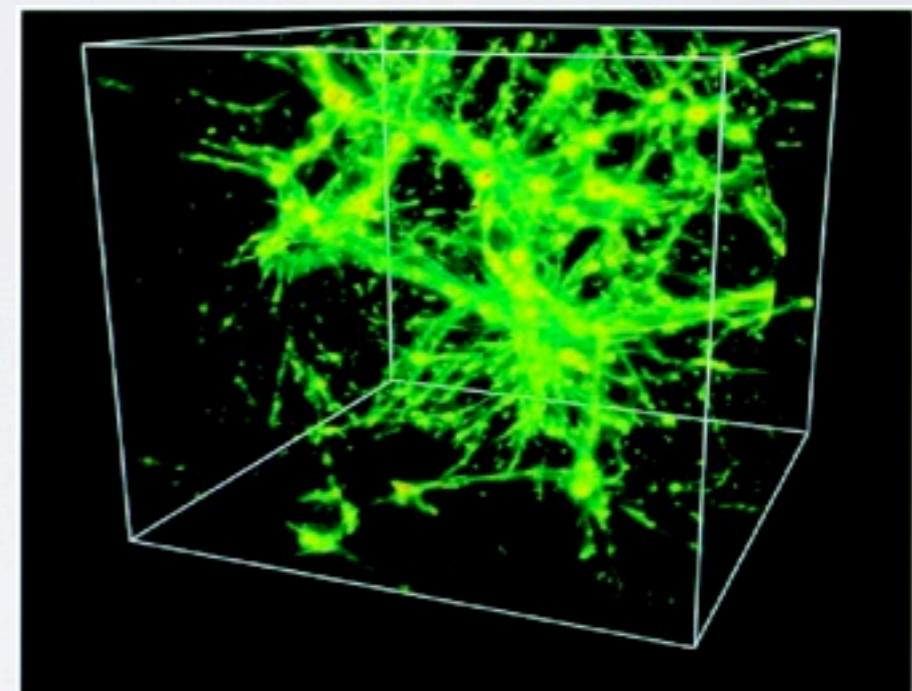
M33/M31 HI STREAM

gaseous tidal stream?

or cosmic infall?

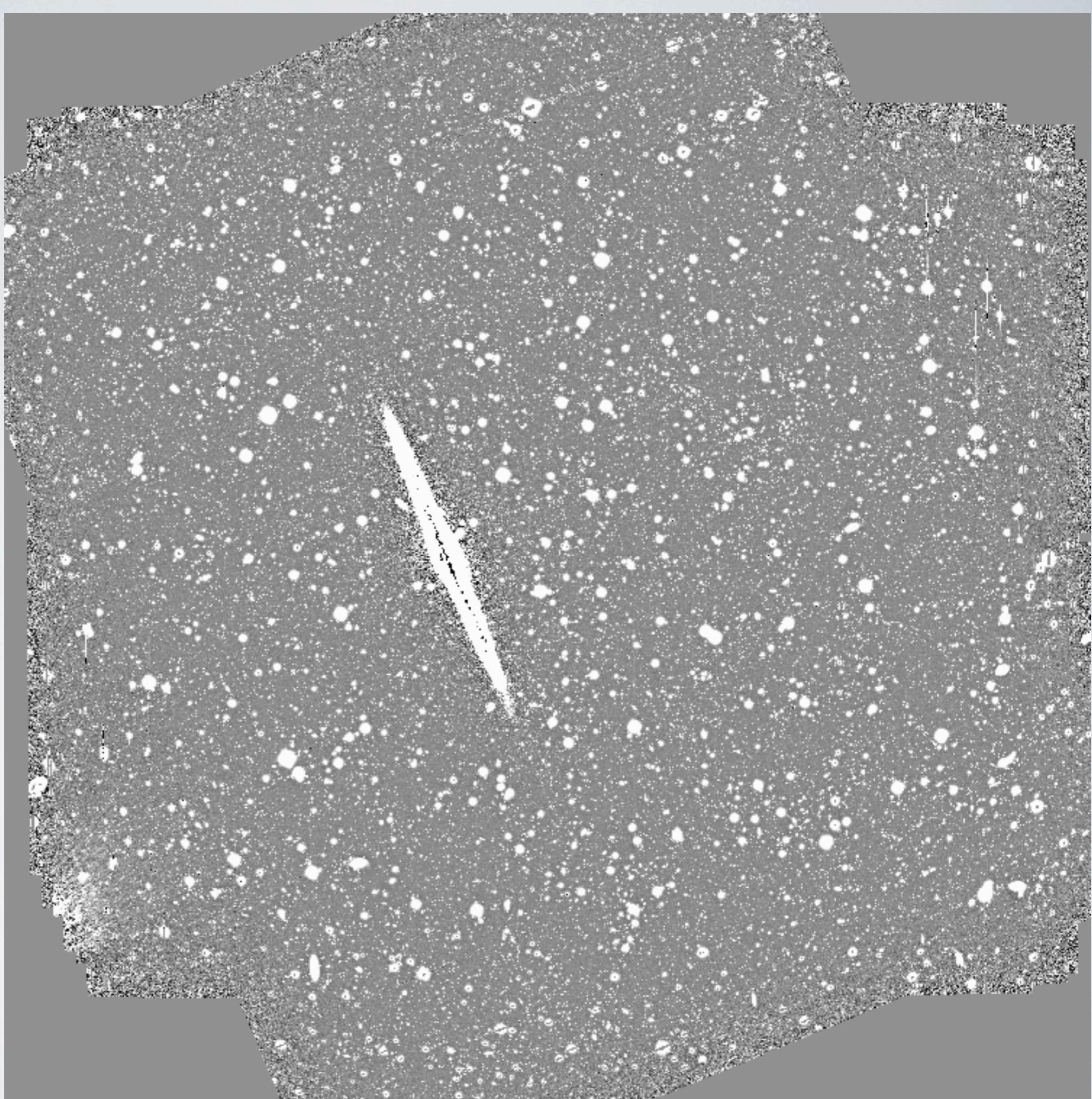


Braun & Thilker (2004)



N89I

- 4 nights
Nov 2008
- 0.4" to 0.8"
seeing!

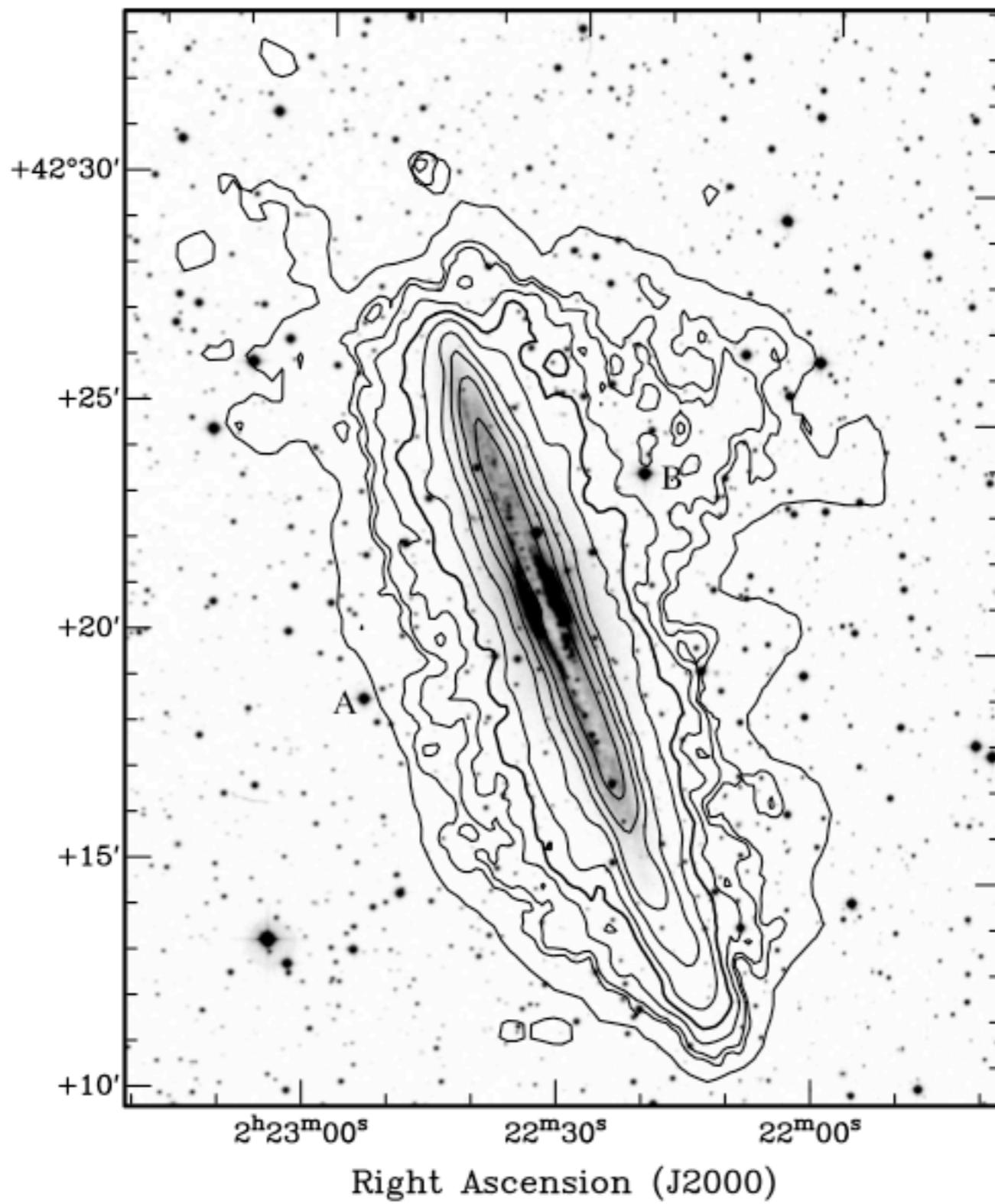


BEYOND M31 ...

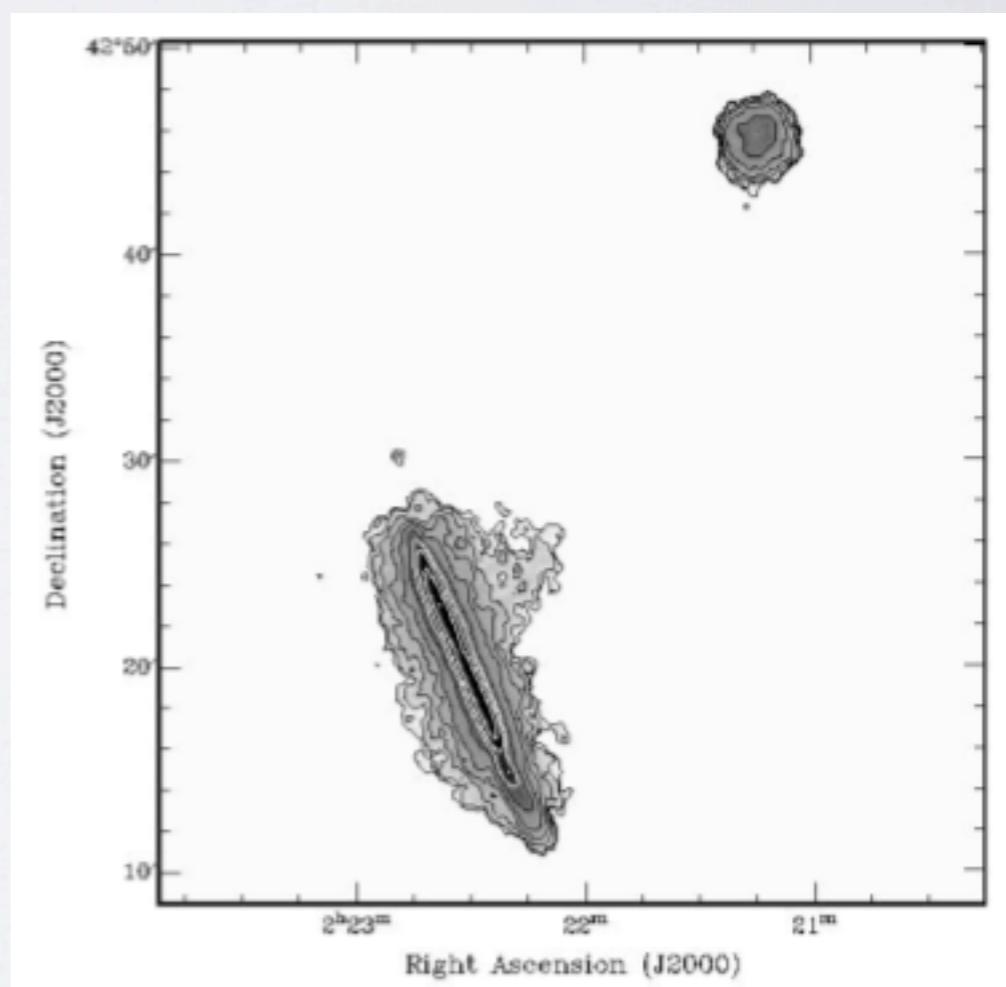
1022

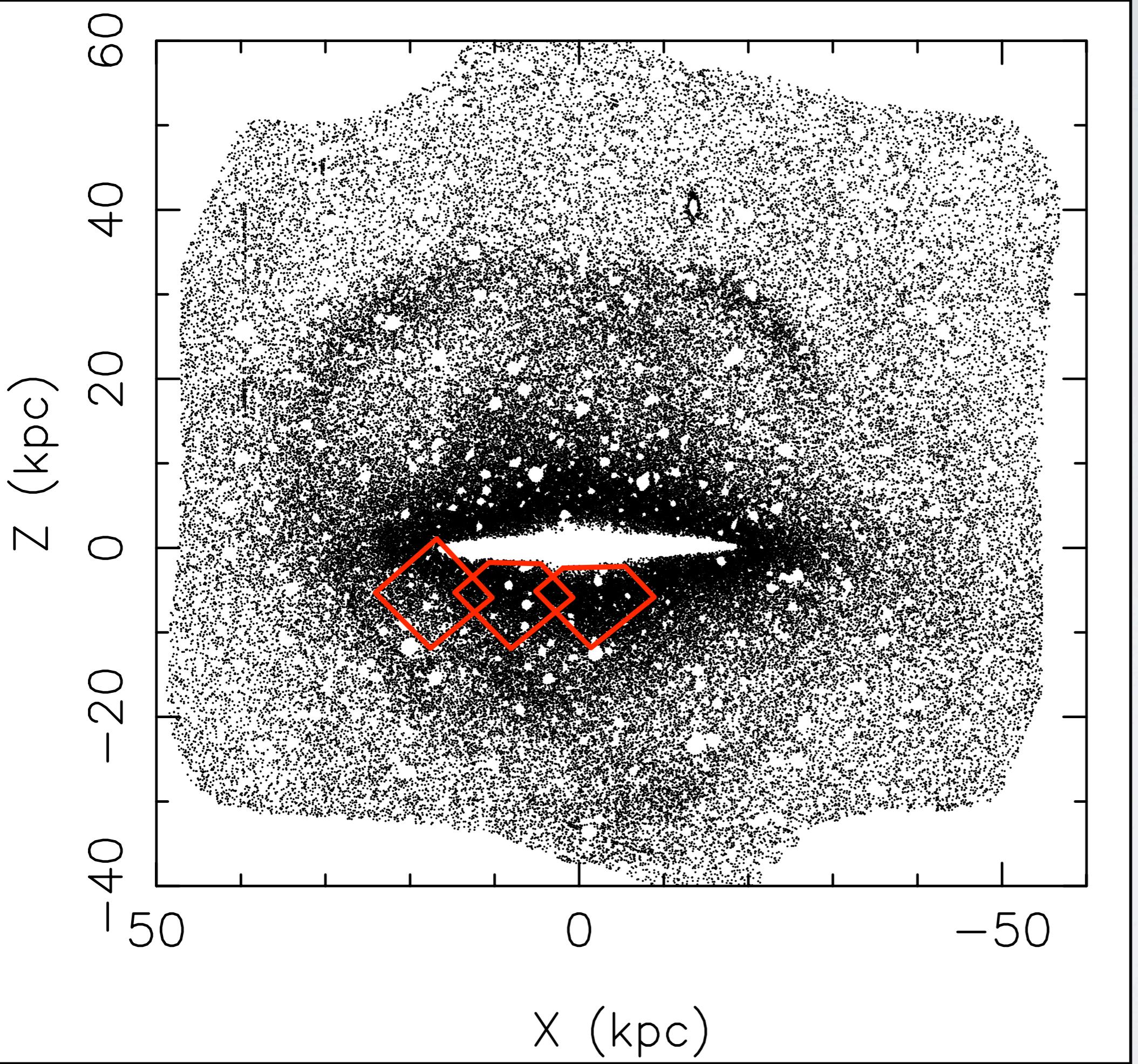
OOSTERLOO, FRATERNALI, & SANCISI

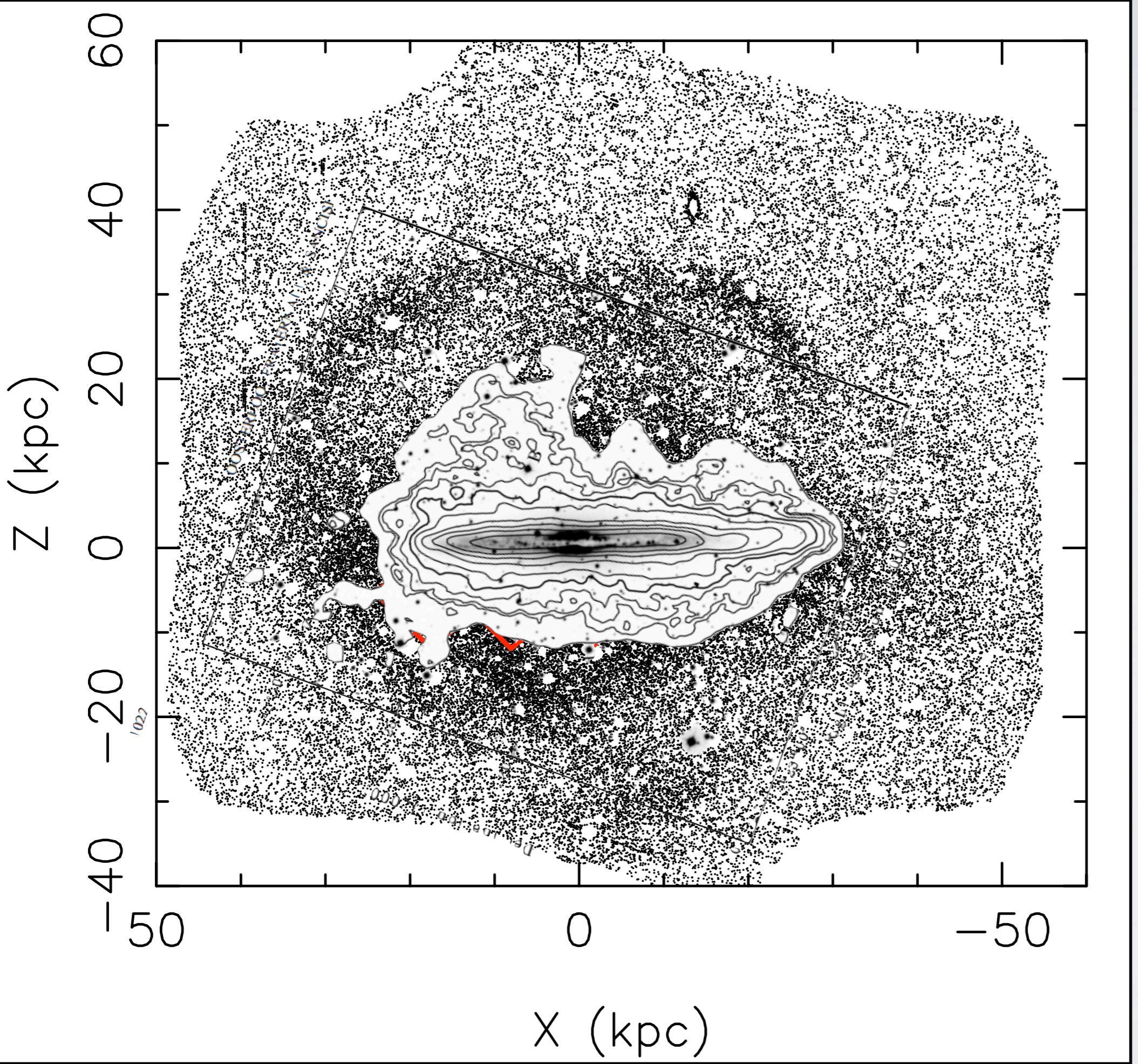
Declination (J2000)

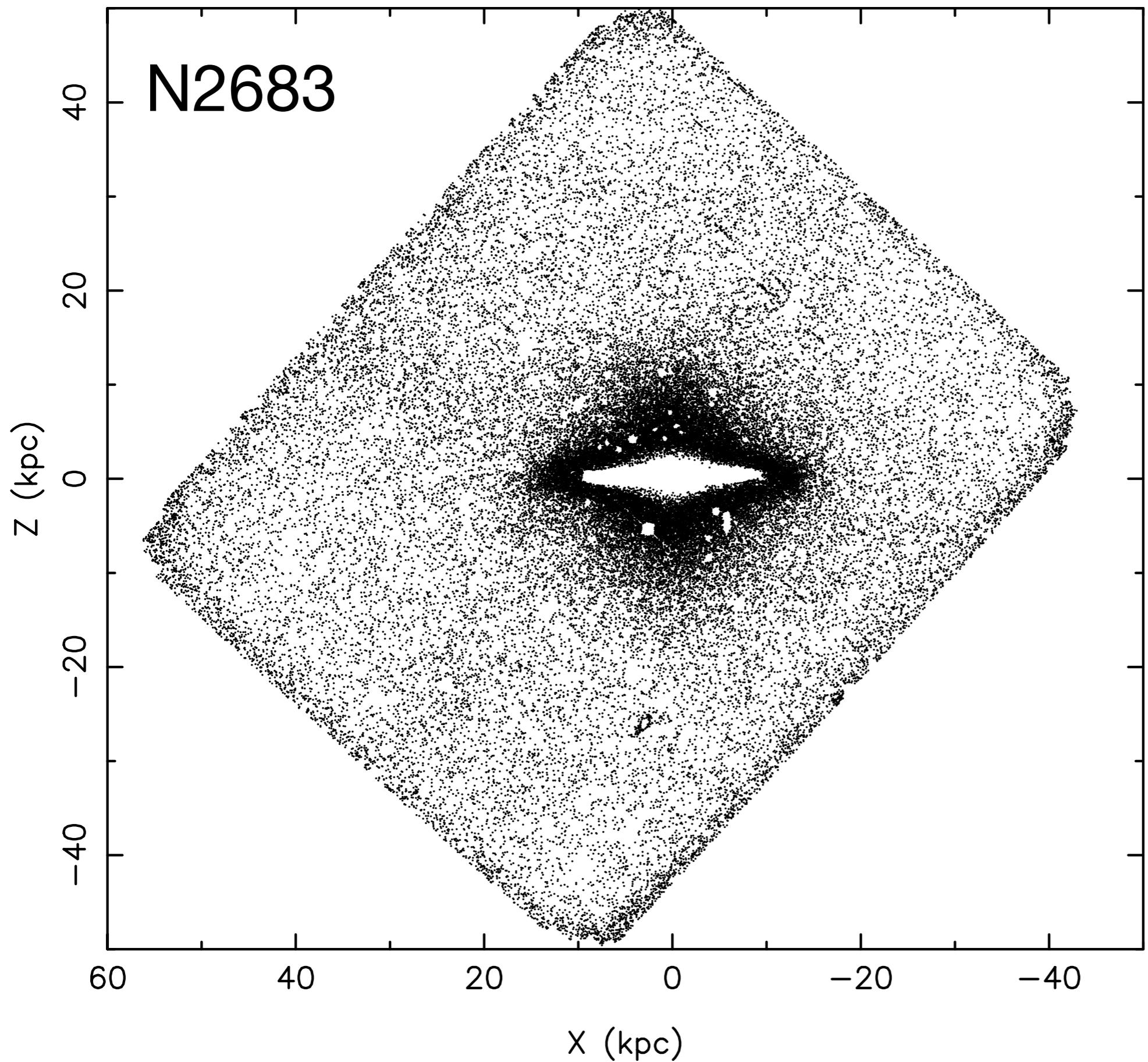


NGC 891





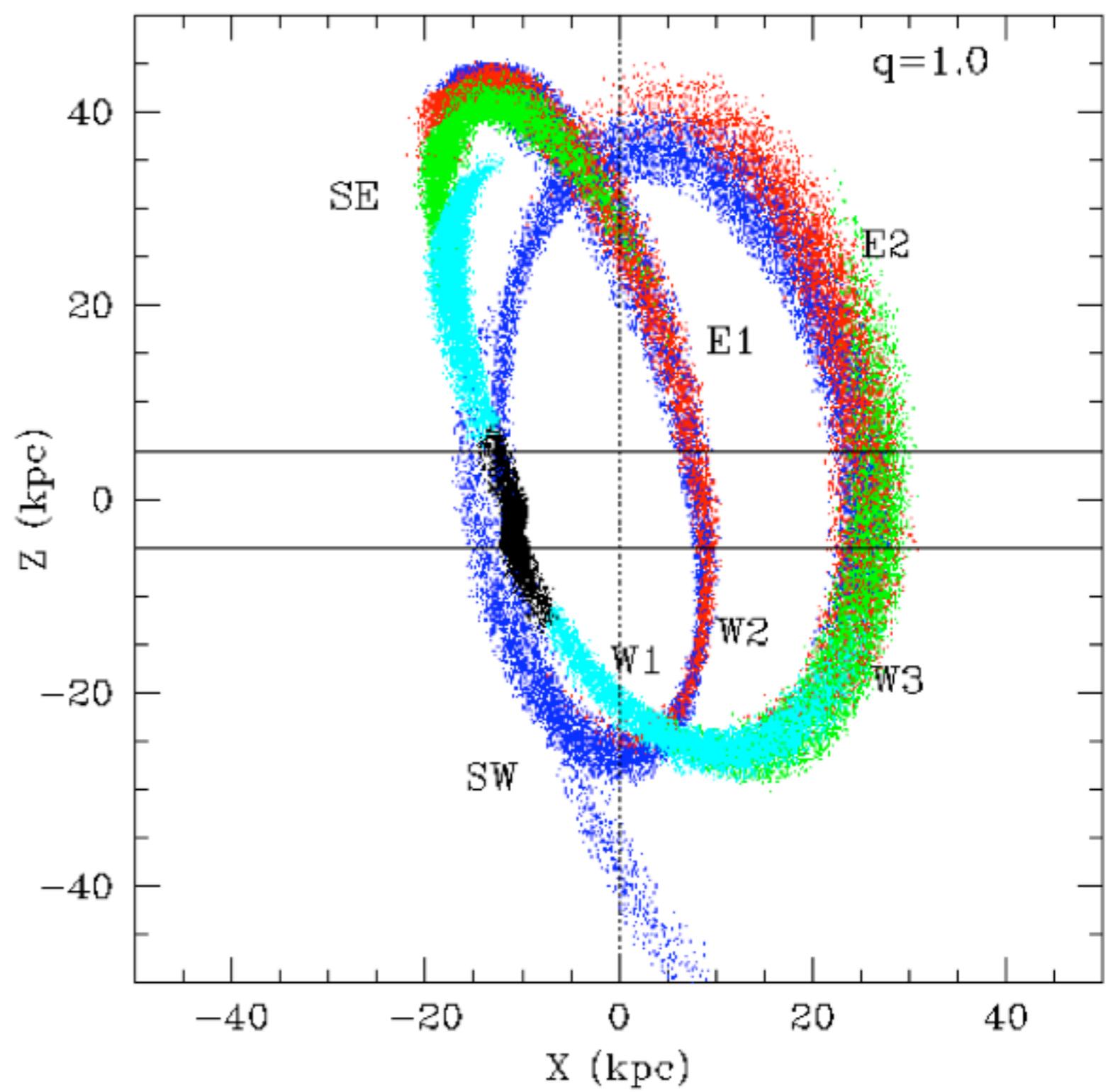




Current
sample:

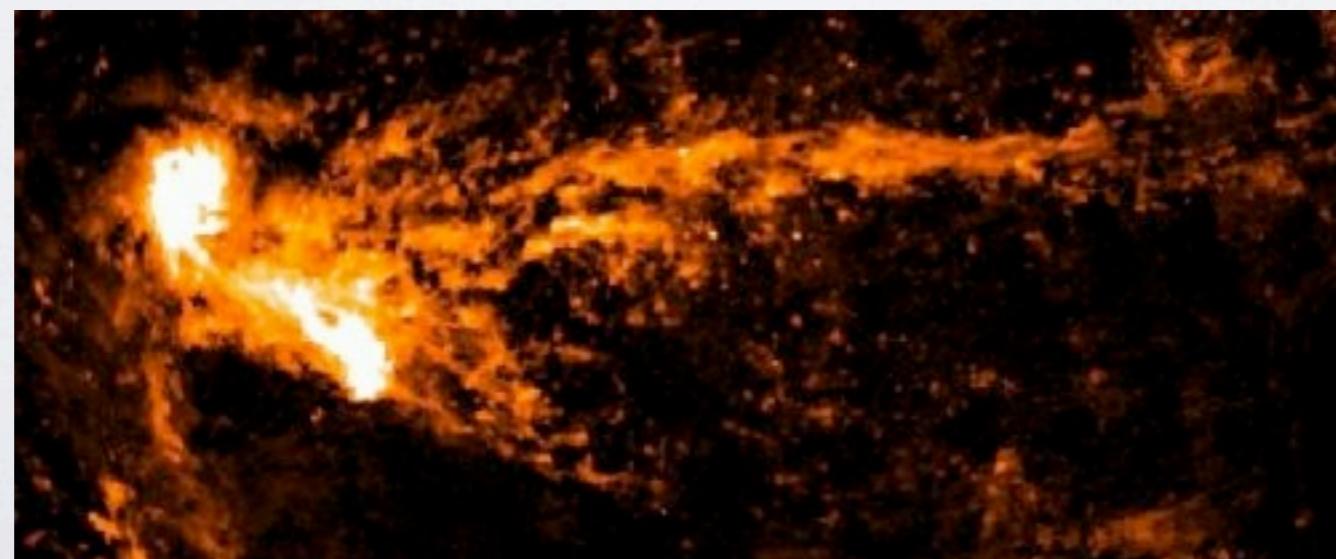
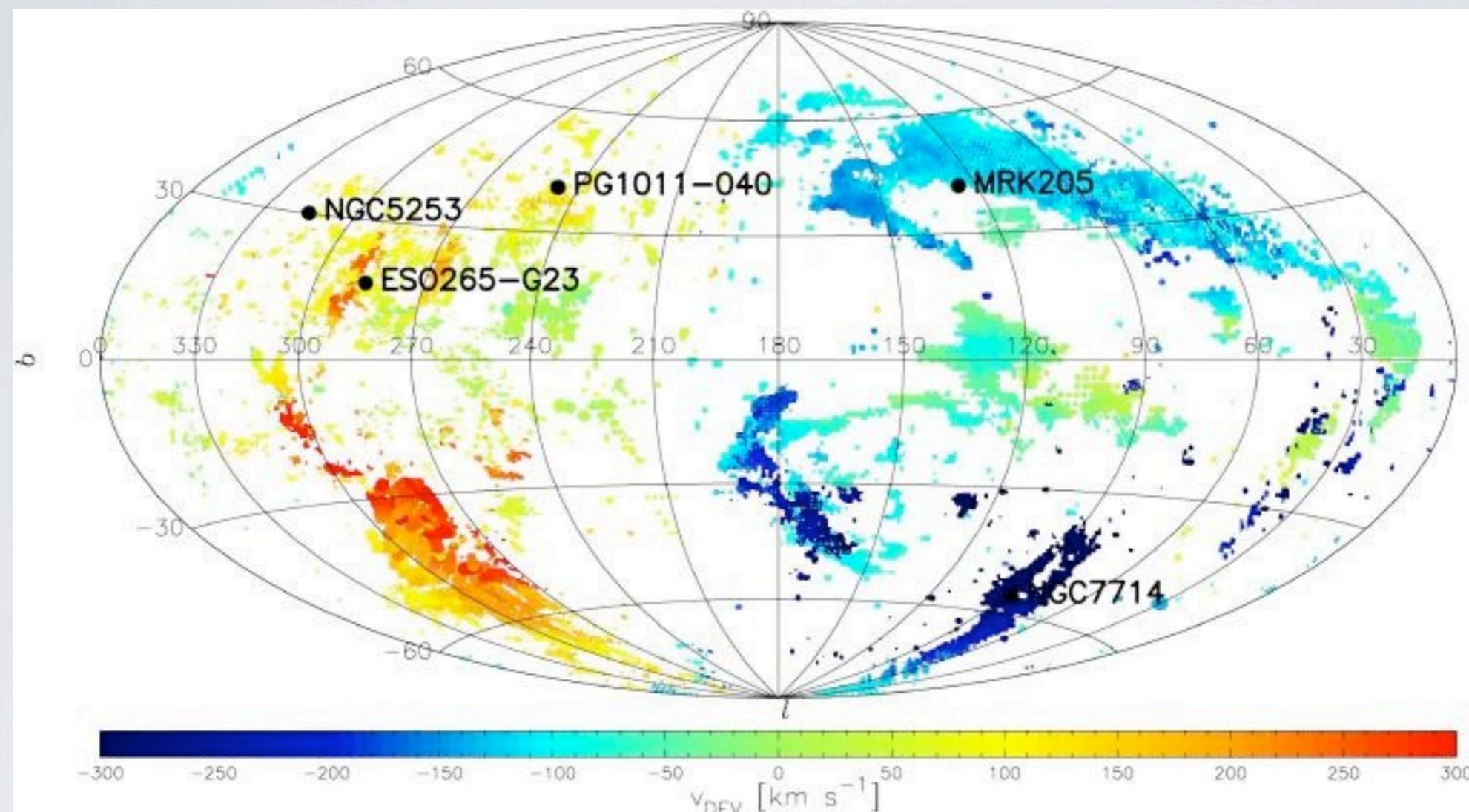
M31
M33
N253
N300
N55
M81
N5128
N4945
N2683
N891

NGC 5907



Martinez-Delgado, Peñarrubia et al. 2008

GALACTIC HI SUBSTRUCTURES



SMALLEST SATELLITES

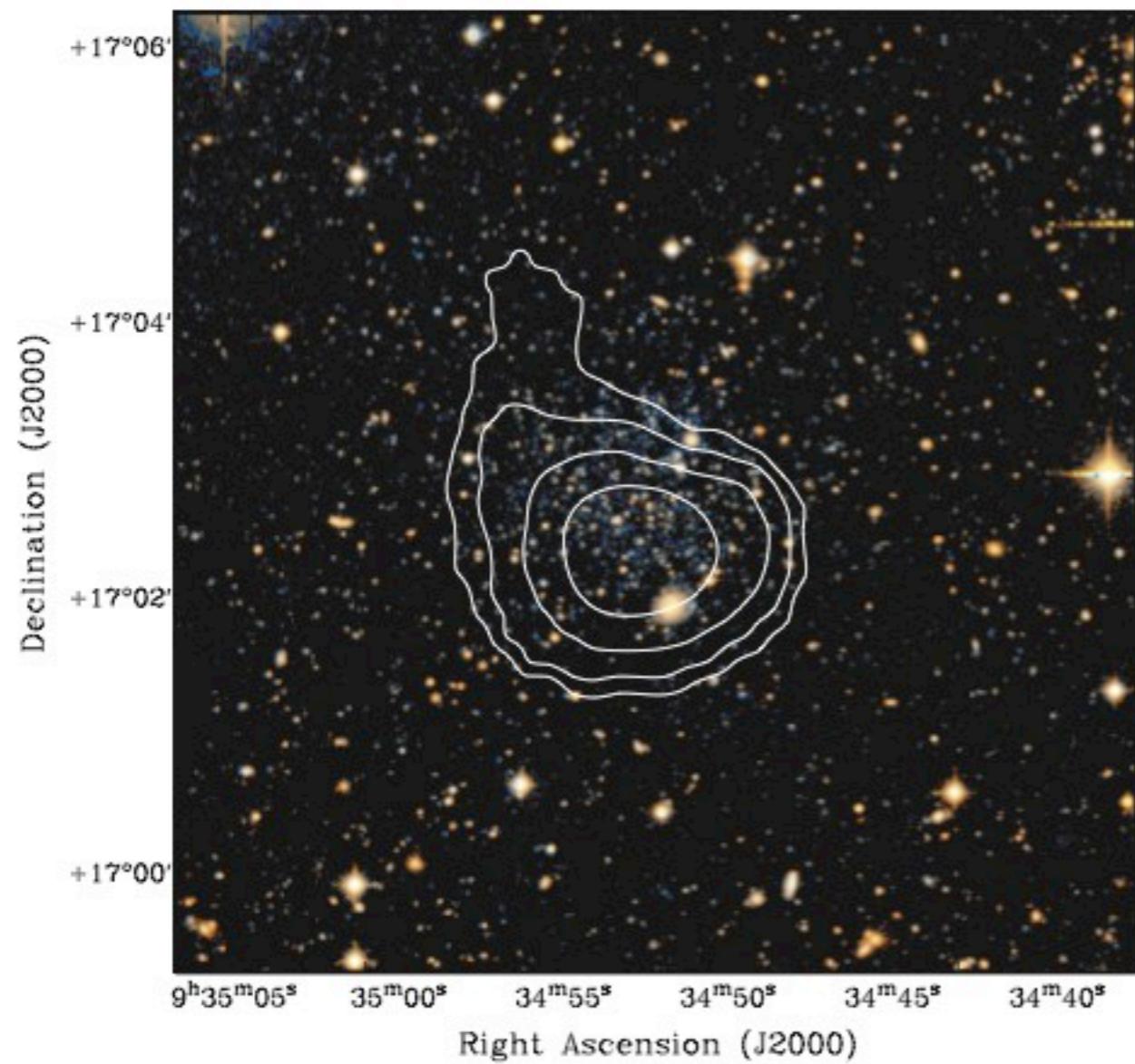


Figure 1. Colour image of Leo T from the Isaac Newton Telescope Wide Field Camera *g*- and *r*-band data with GMRT HI contours overlaid. The column density contours at $2, 5, 10$ and $20 \times 10^{19} \text{ cm}^{-2}$, and the beam size is $39 \times 47 \text{ arcsec}^2$.

Leo T & sisters...?

Irwin et al. (2007)
Ryan-Weber et al. (2008)

420kpc
 $3 \times 10^5 M_\odot$ HI gas

SOME (OBVIOUS!) PROJECTS... WITH SKA PRECURSORS

- Survey of gaseous accretions onto disks (preparatory optical studies?)
- Survey of minor interactions (a la M31 vs. M33; NGC 891)
- Use gas-rich satellites as statistical dark matter tracers?
- Correlations between direction of accreting gas and L_z of disk
- Put Sancisi's Rule on firm statistical footing
- Detection of smallest gas-rich galaxies in Local Group (in conjunction with Skymapper/LSST). Tie-in to cosmic web?
- Detection of older gaseous streams orbiting Milky Way (Sagittarius? Fornax? ...currently may not be looking in the right place...)