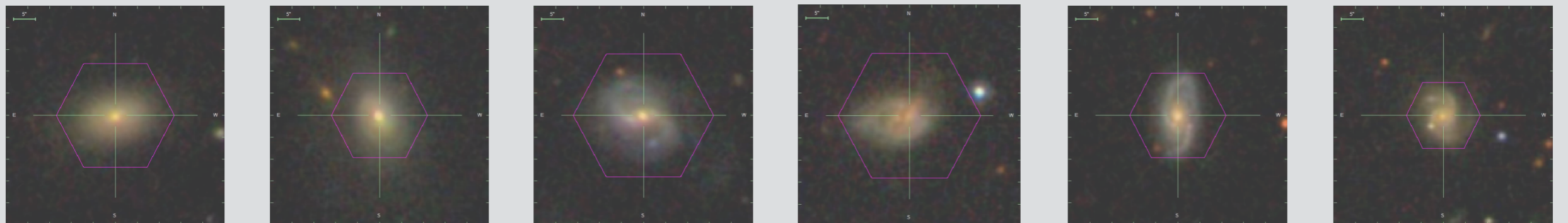
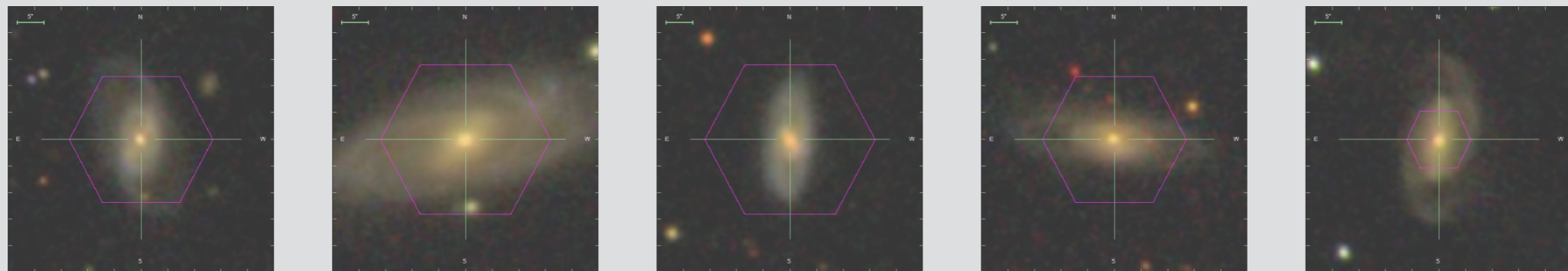


Asymmetric kinematics features within galaxies at $z=0.06$ as revealed by MaNGA

Barbara Mazzilli-Ciraulo, supervised by Anne-Laure Melchior and Françoise Combes



Detecting double-peaked spectral features

RCSED 

Value-added catalog
800,299 galaxies
from SDSS DR7



Double peak catalog

Maschmann et al. (in prep)

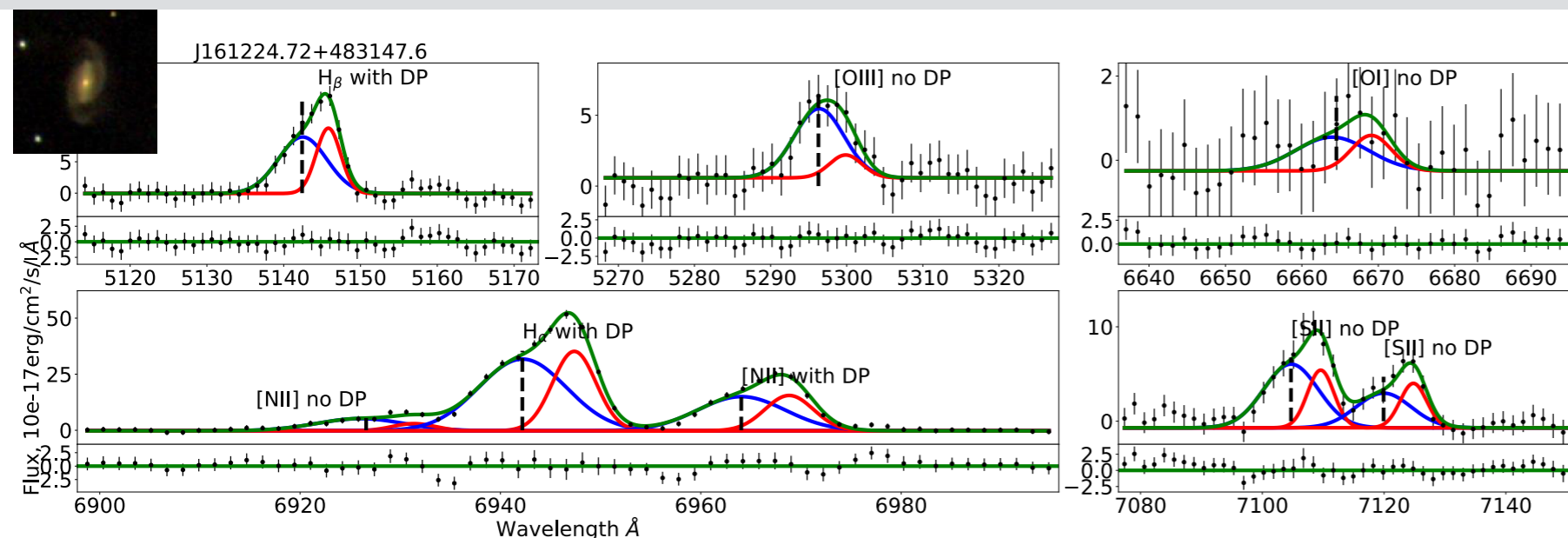
5289 galaxies

0.6%

Maschmann & Melchior, arXiv:1906.05629



40% of S0
2 different components
Isolated
Ongoing star formation

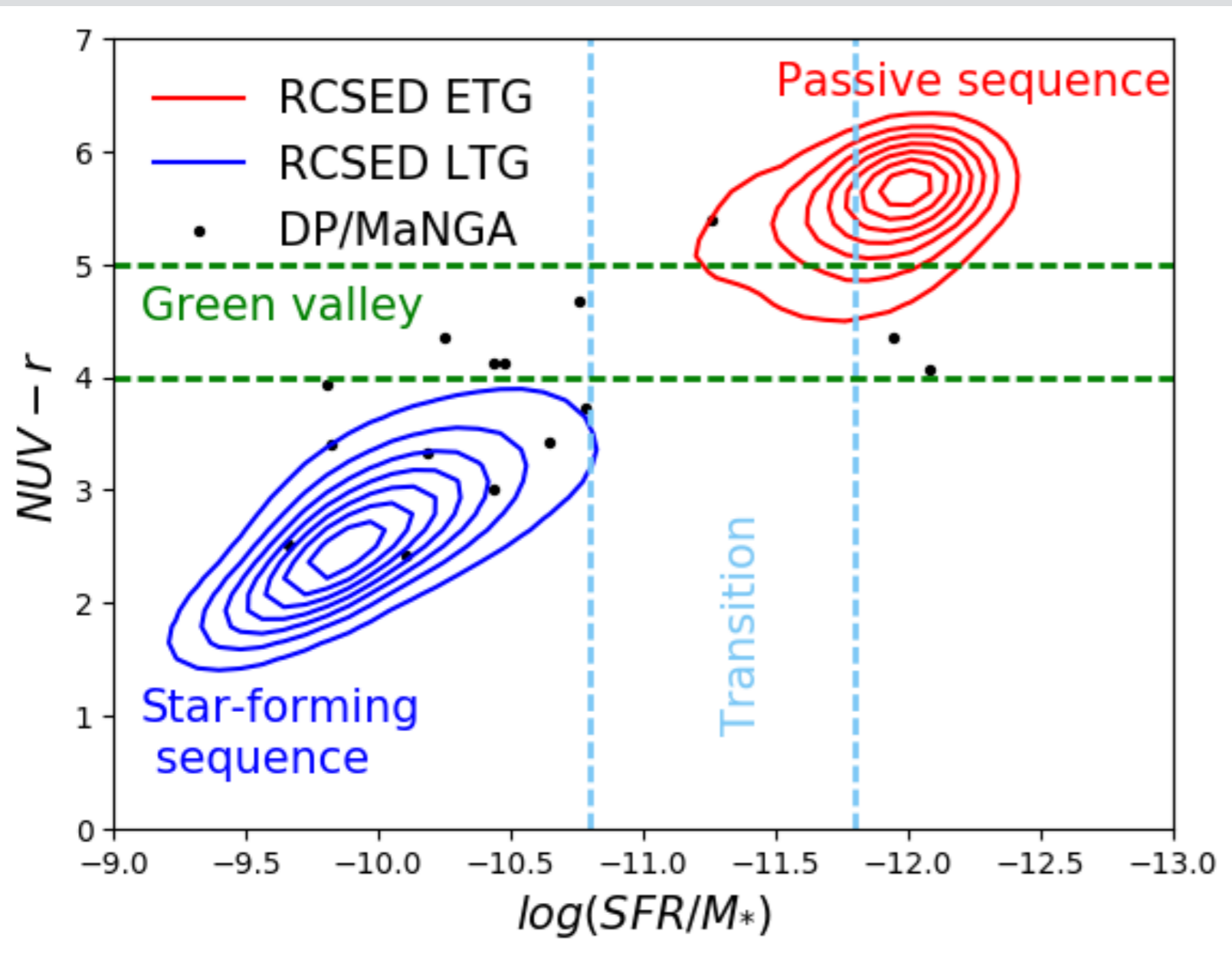


3-arcsec SDSS fiber integrated spectrum

Double peaks: a tool to investigate quenching?

→ S0 lead to ellipticals
Ellipticals lie on the red sequence

→ S0 as a result of mergers
(Fraser-McKelvie et al. (2018)
and Eliche-Moral et al. (2018))



Passive sequence

Old/retired galaxies

The Green Valley:

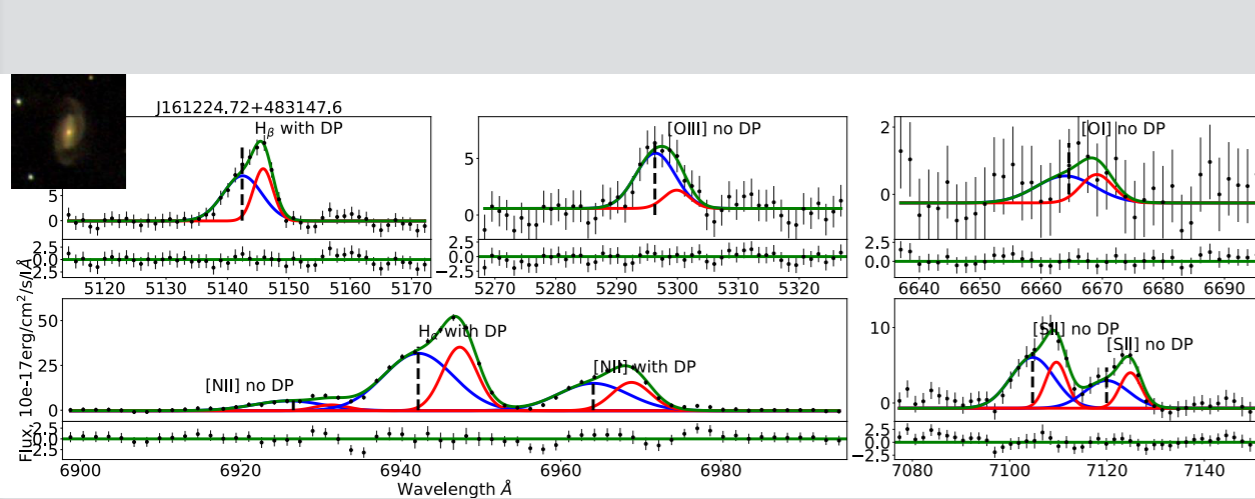
Low levels of ongoing
star formation

Star-forming sequence

Young galaxies
Active star formation

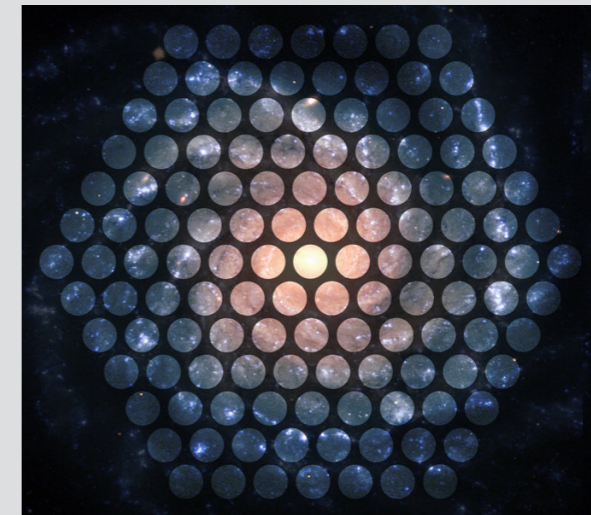
Double peak catalog: 5289 galaxies

Maschmann & al. (in prep.)



MaNGA survey (DR14): 2618 galaxies

Bundy & al. (2015)

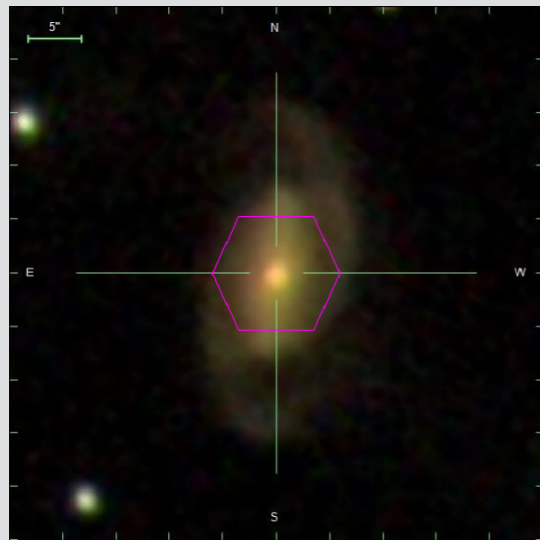


<https://www.mpia.de>

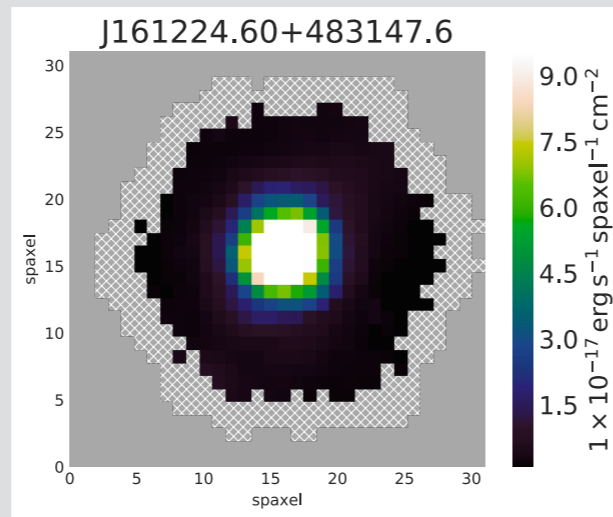
17 galaxies ($\langle z \rangle \sim 0.06$)

**Second component?
Minor-merger?**

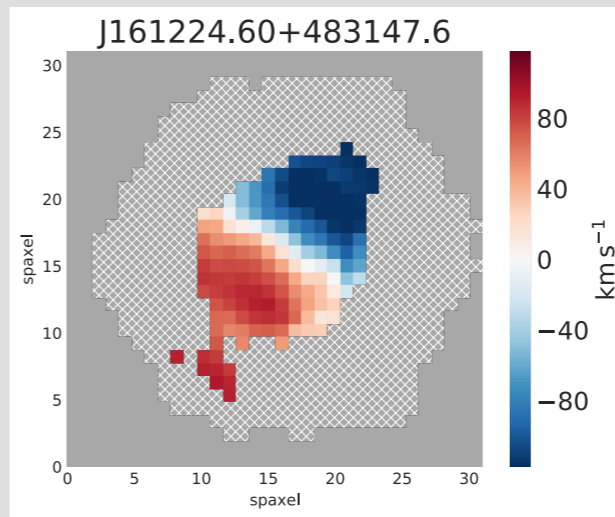
Maps as produced by the Data Analysis Pipeline



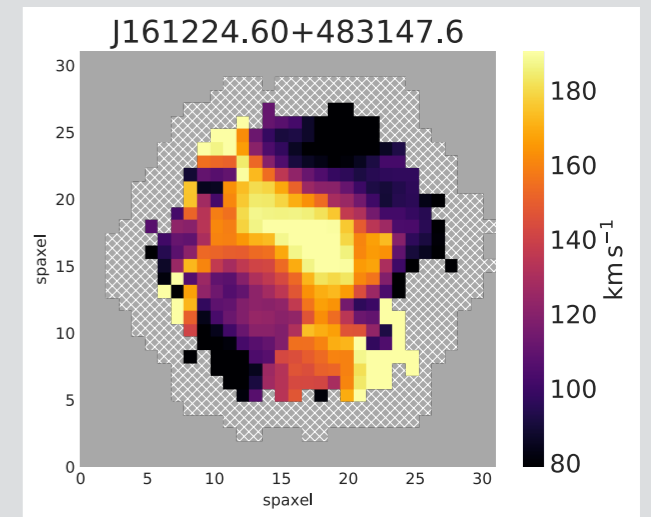
J 161224.6+483148



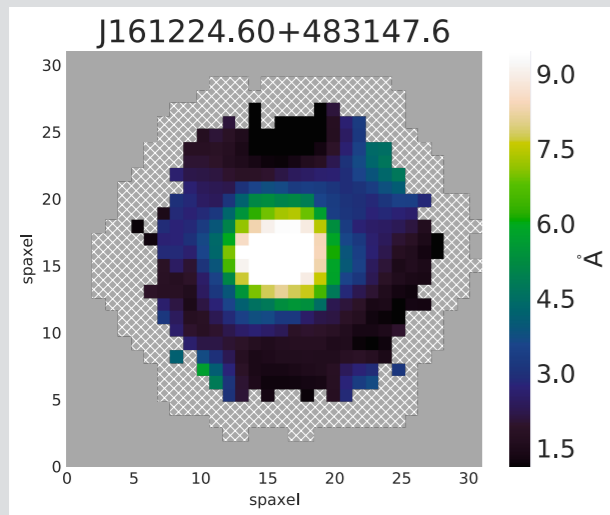
H α flux



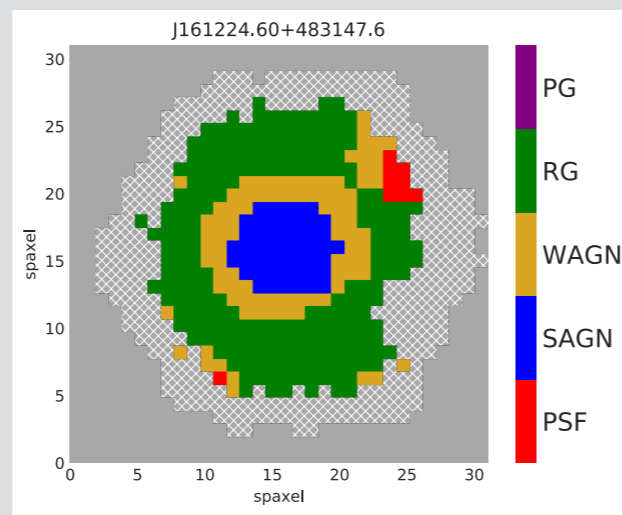
Gas velocity field



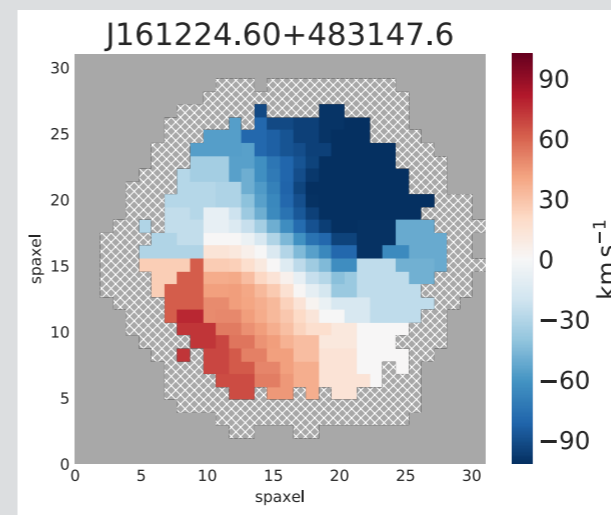
H α sigma



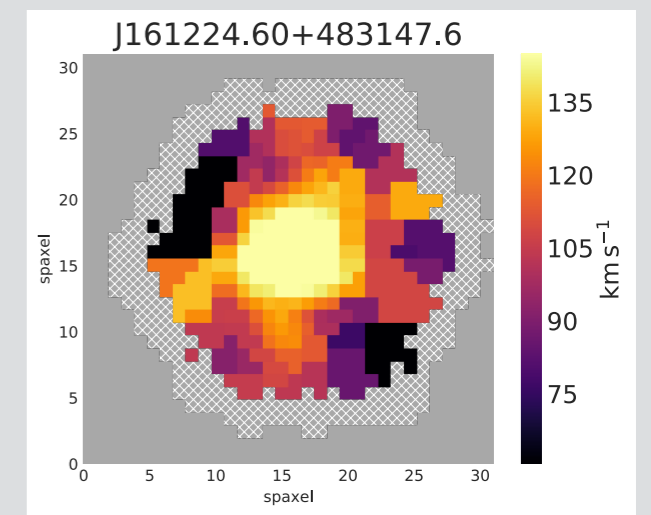
H α equivalent width



WHAN diagram



Stellar velocity field

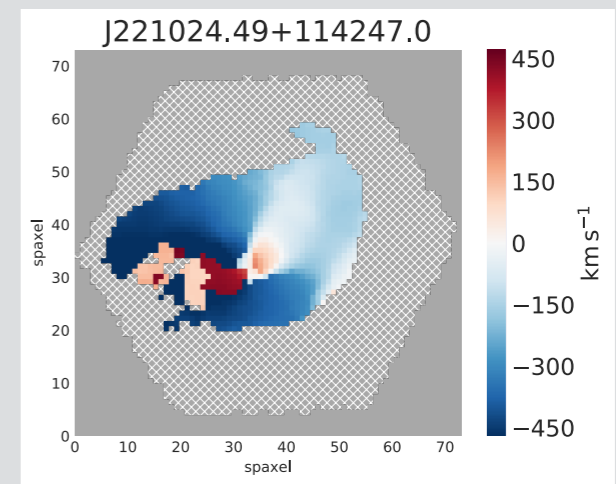
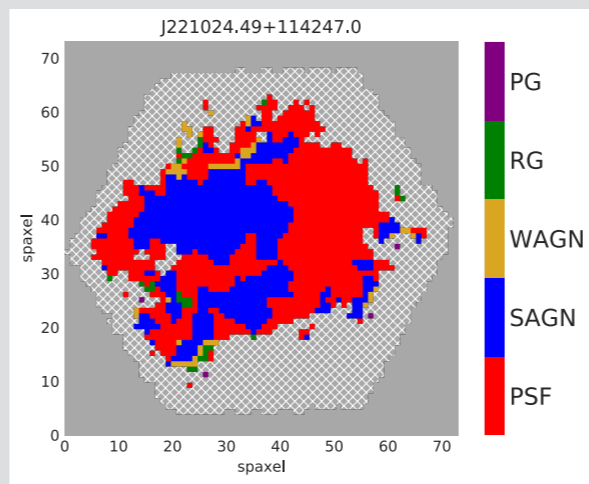
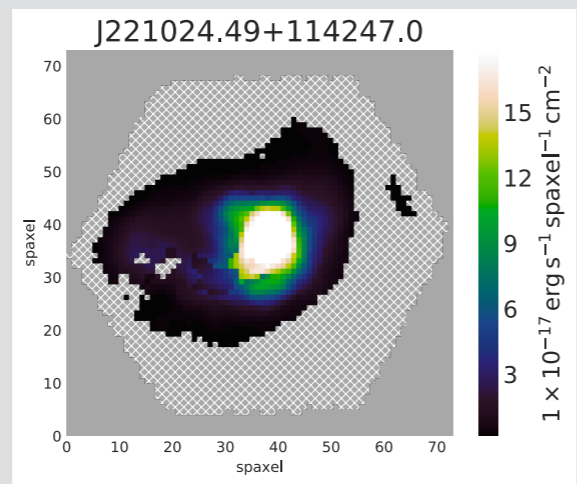
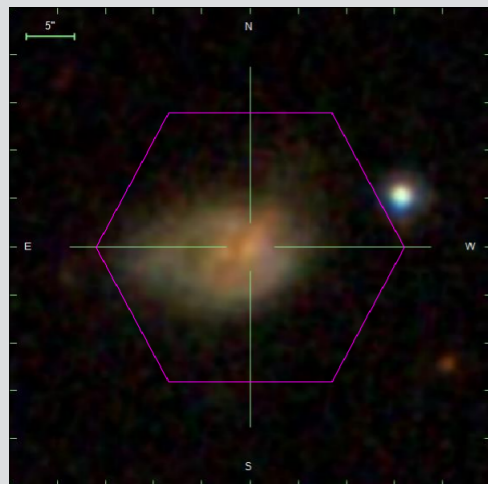
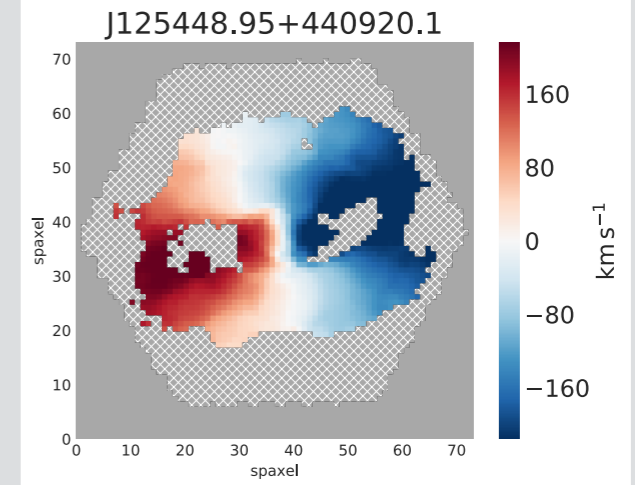
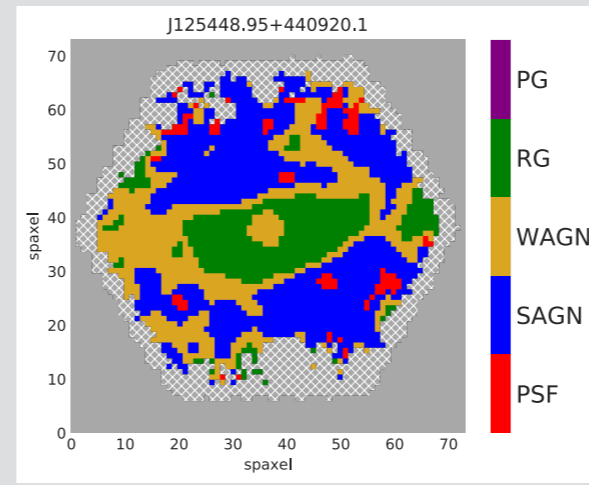
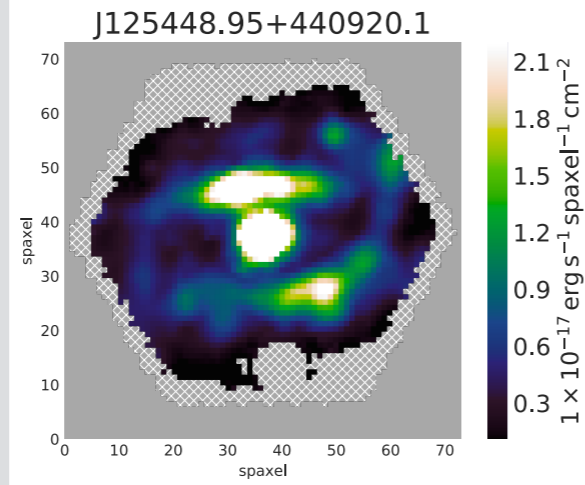
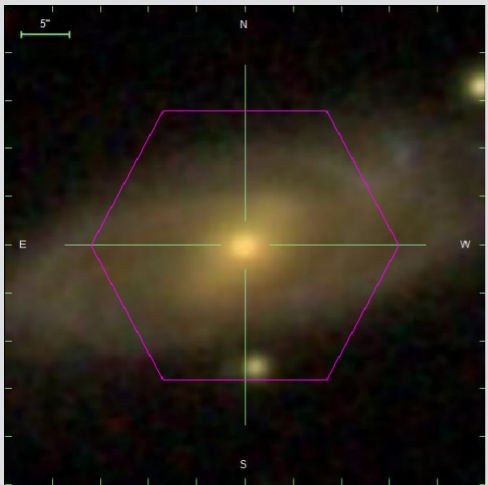
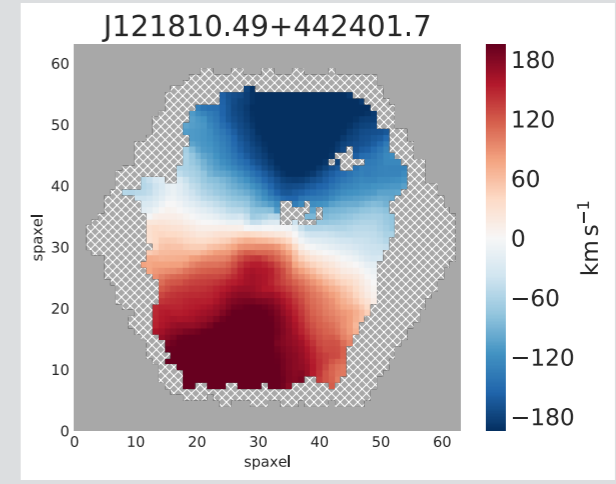
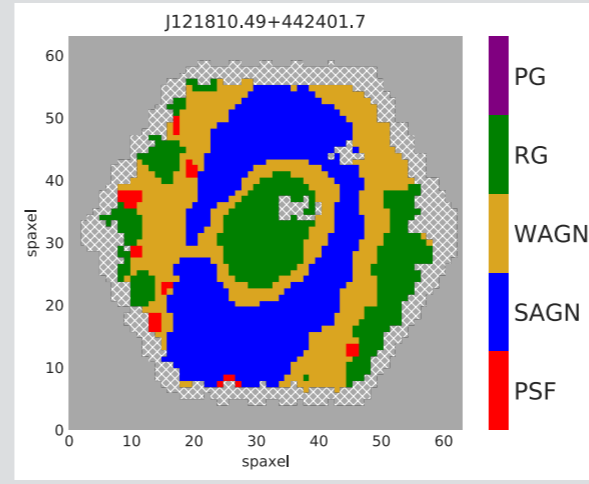
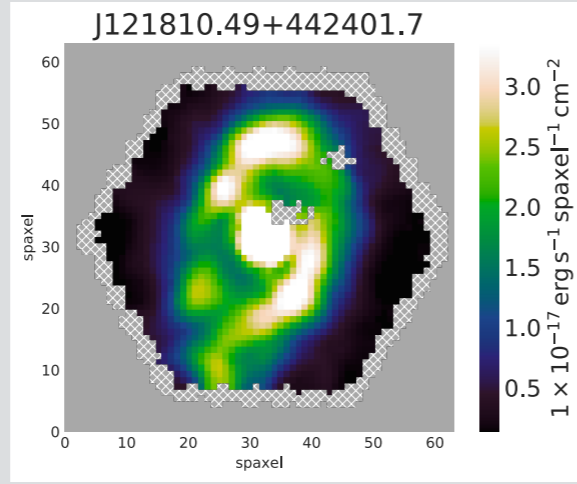
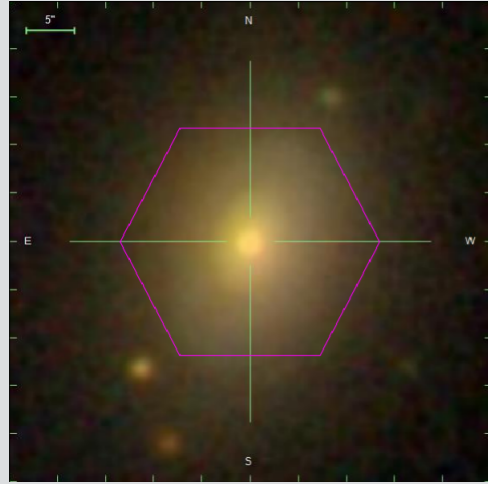


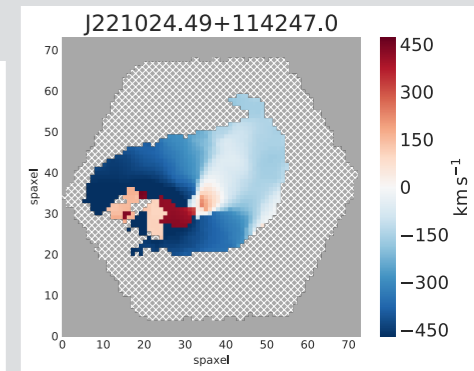
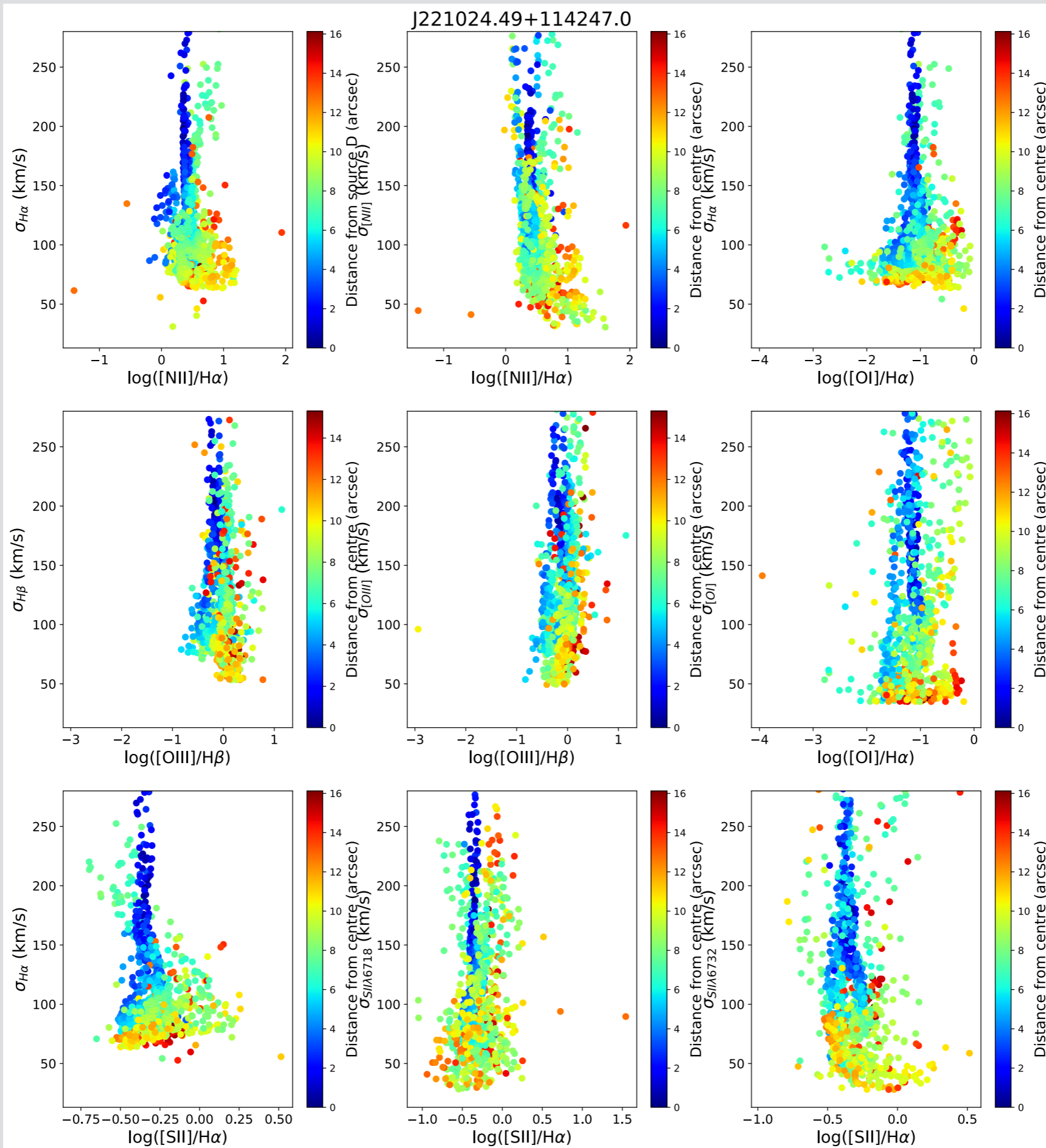
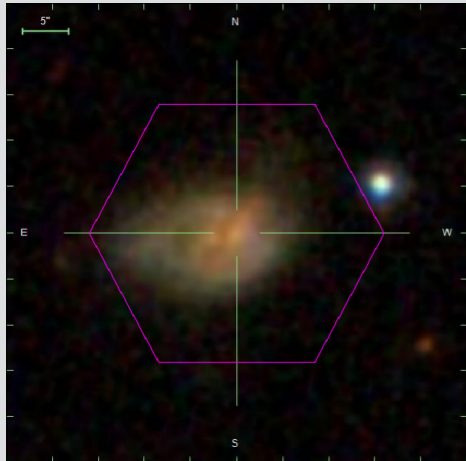
Stellar sigma

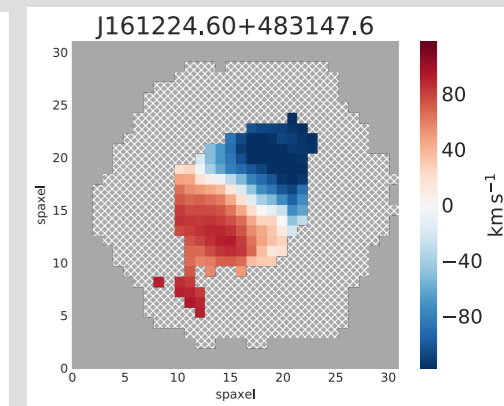
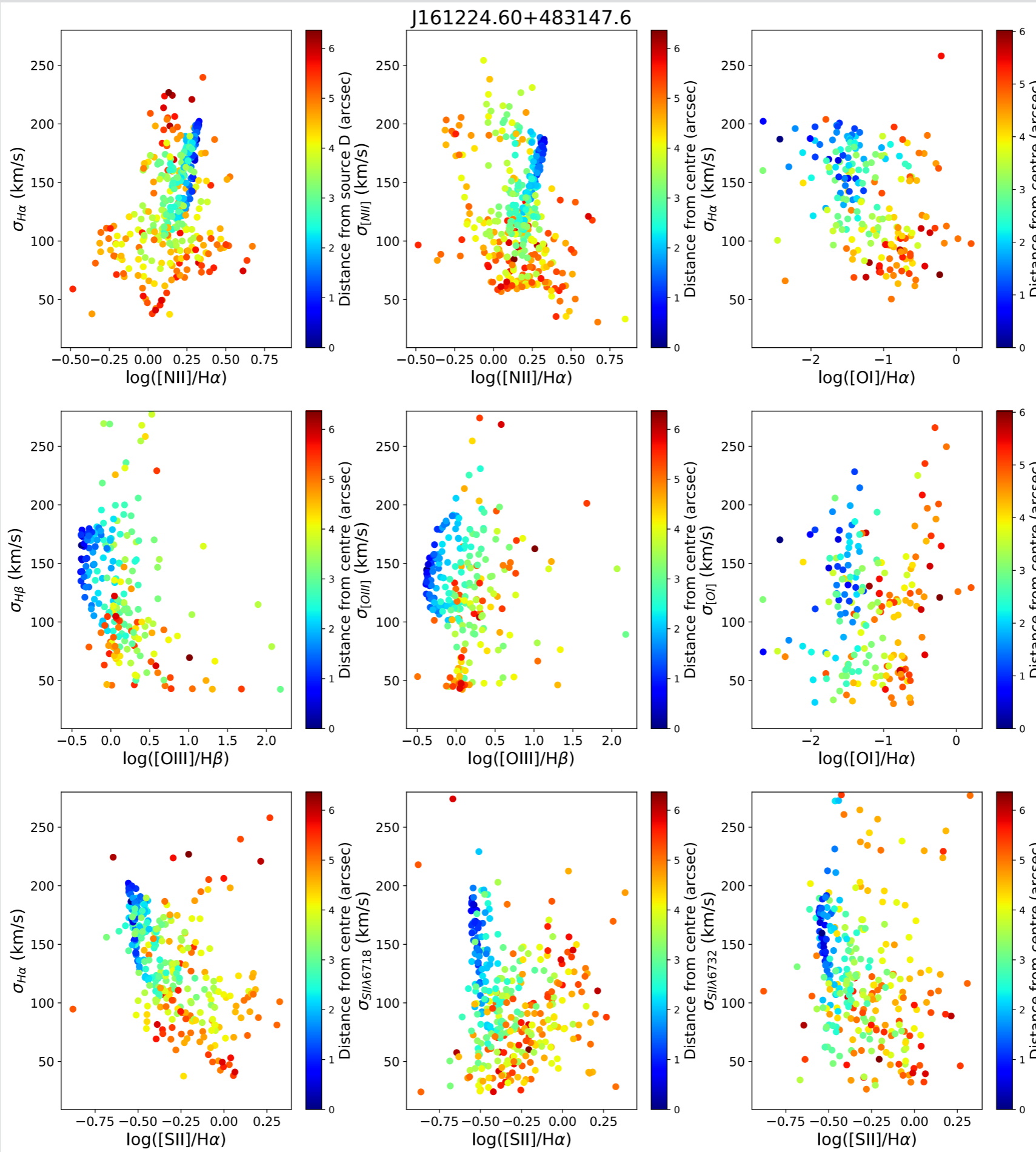
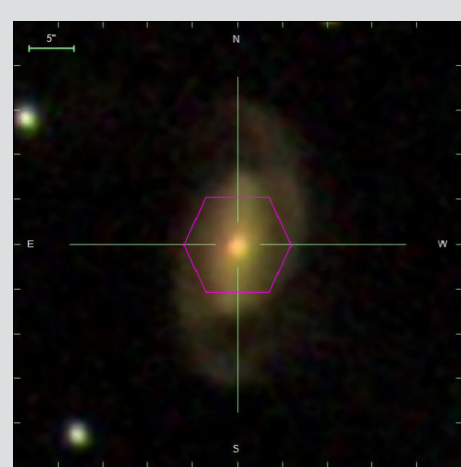
H α flux

WHAN diagram

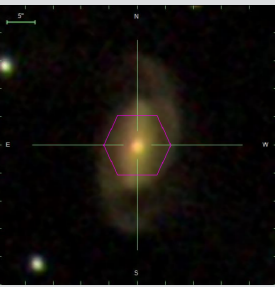
Gas velocity field



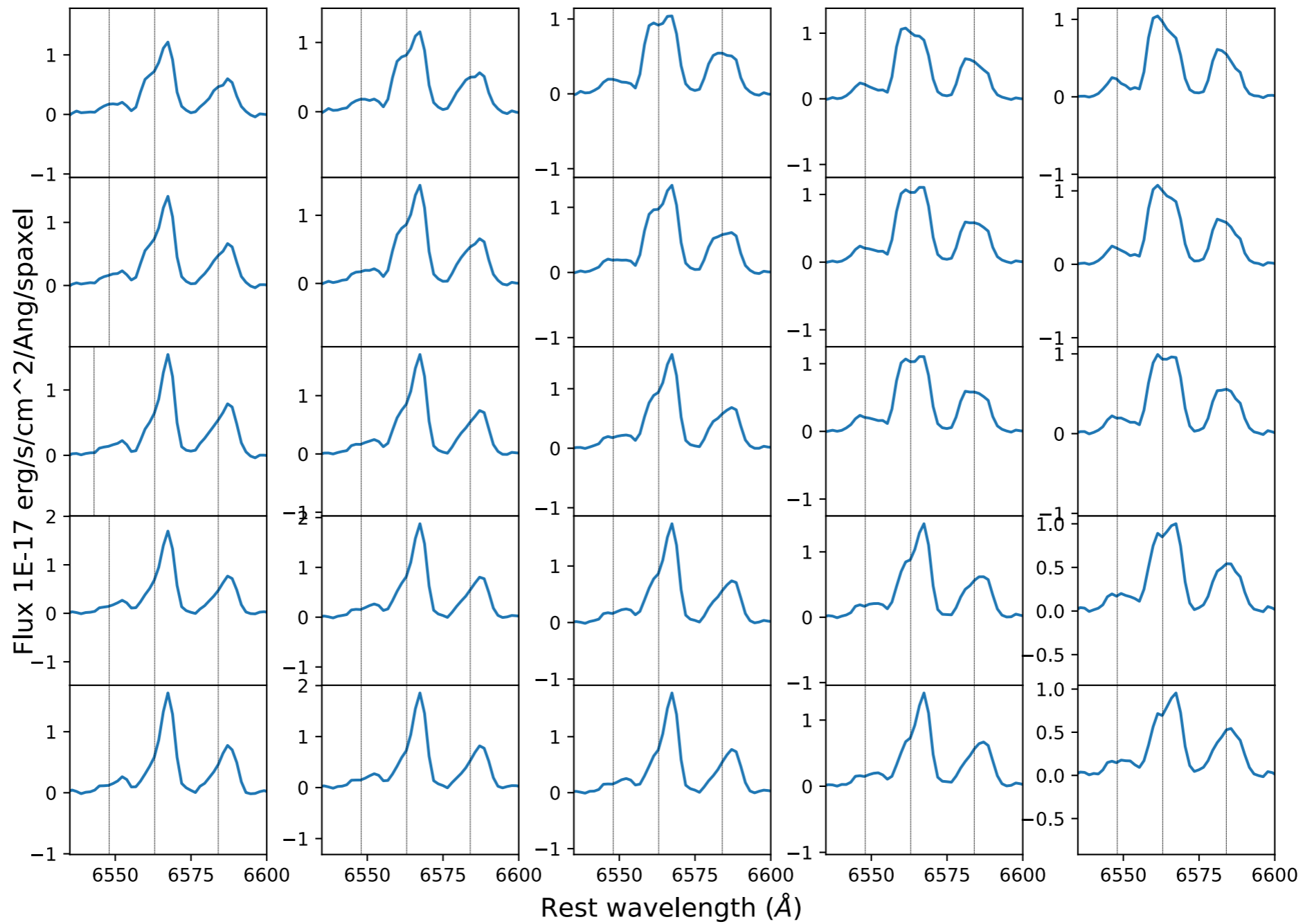




But...



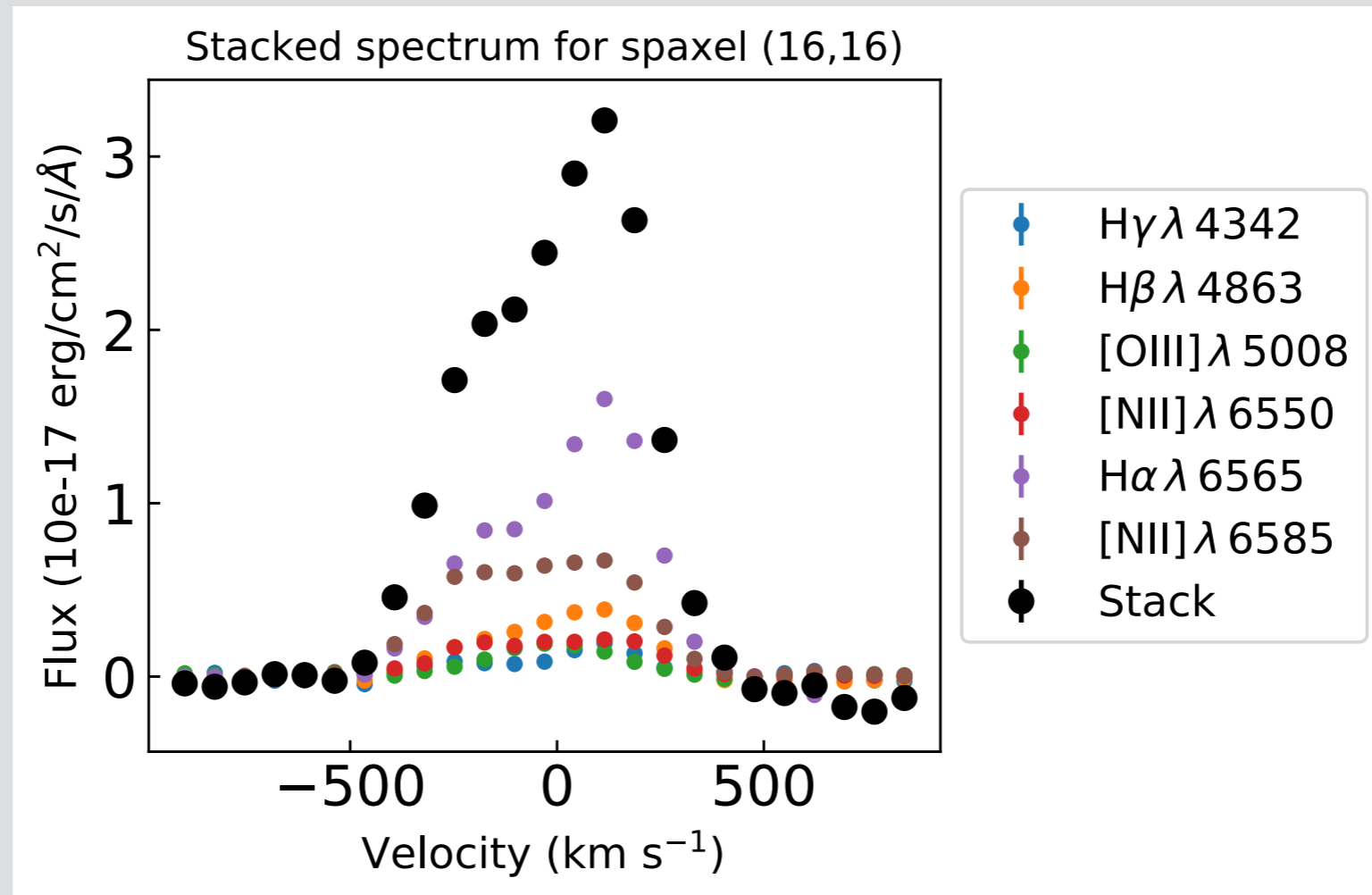
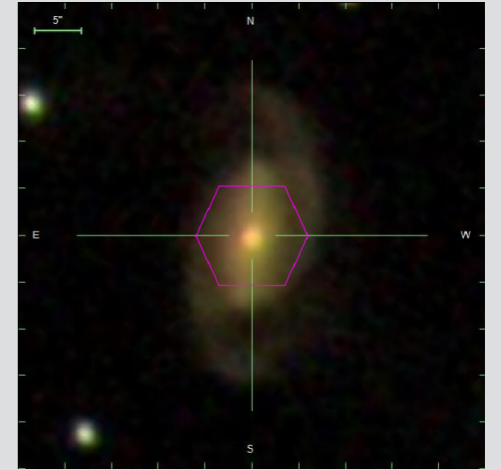
J161224.6+483148



No consideration for the double-peaked features

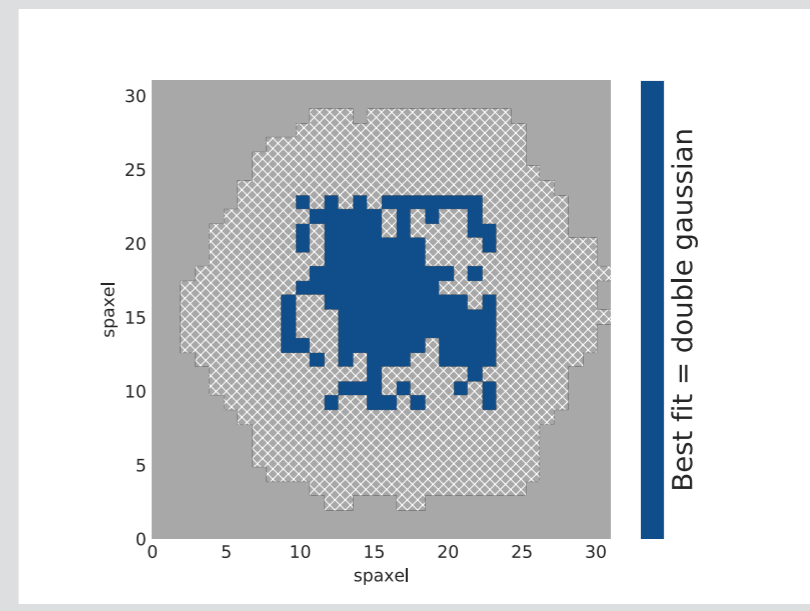
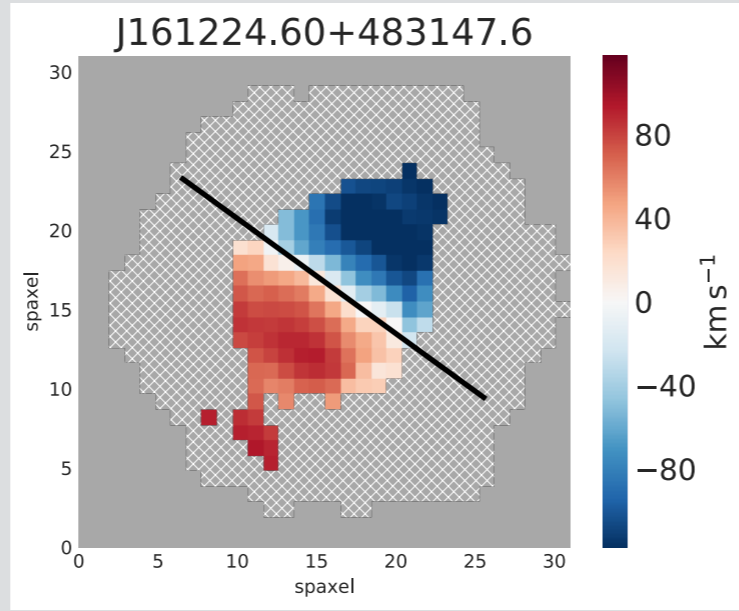
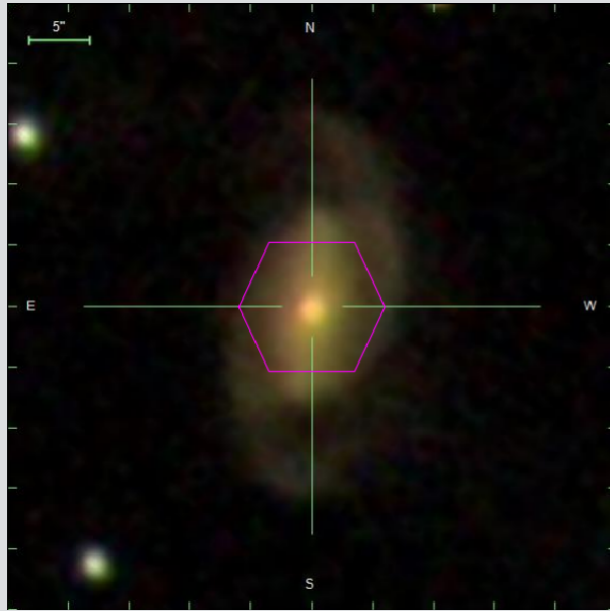
Fitting procedure

Stack all the lines for each spaxel



One single μ , one single σ for all the lines

Fit one single gaussian and a double gaussian



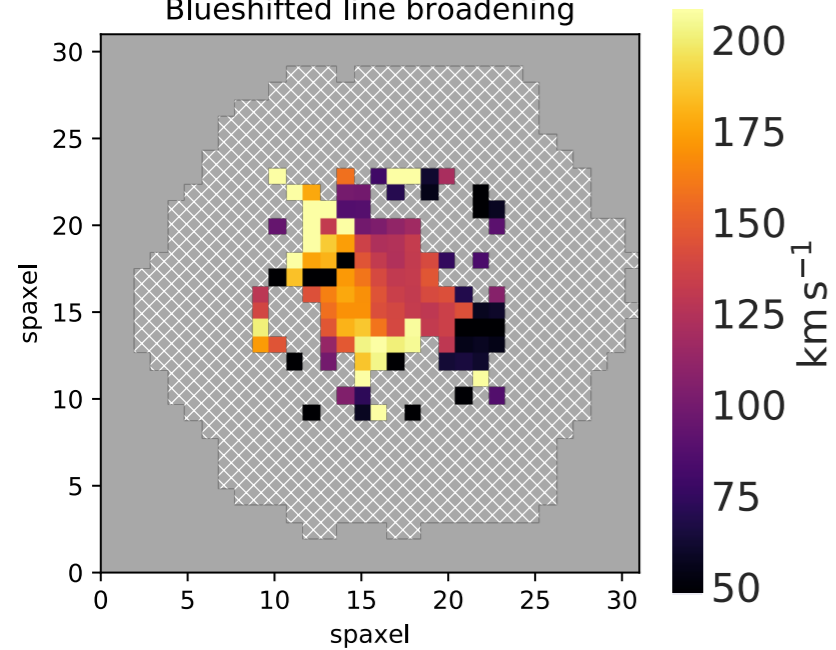
Single component vs. double components

Set of maps:
Strongest component

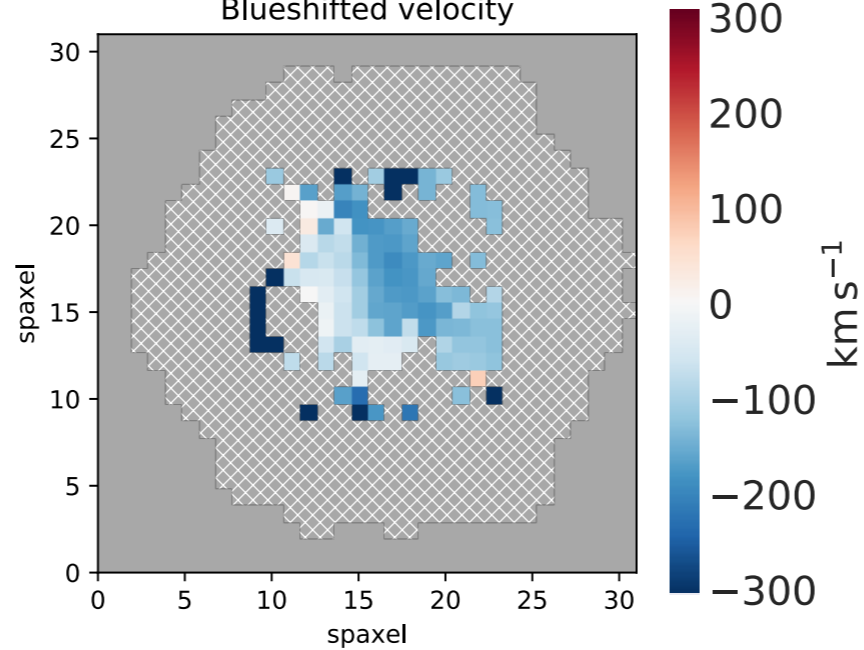
Set of maps:
Secondary component

Distribution and velocity for each

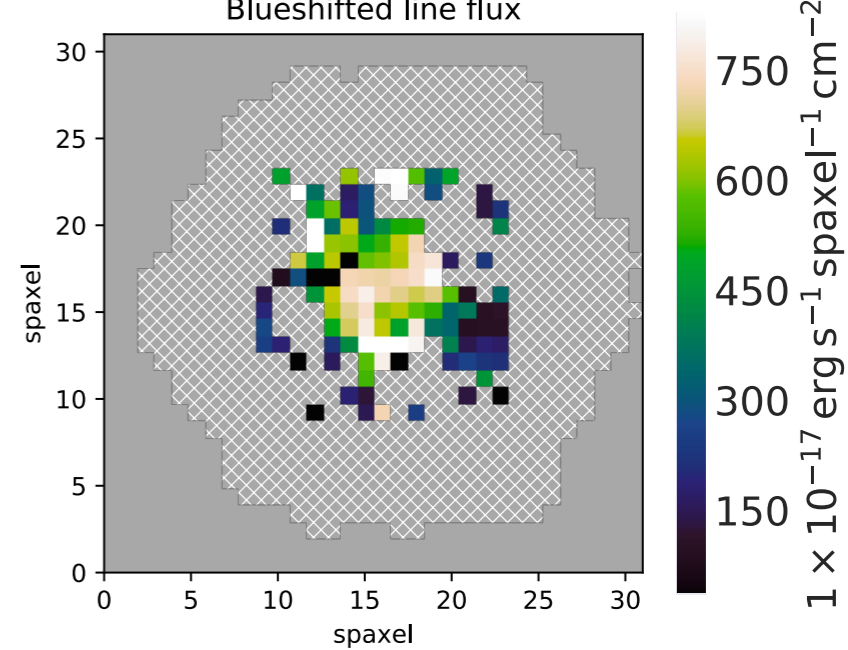
Blueshifted line broadening



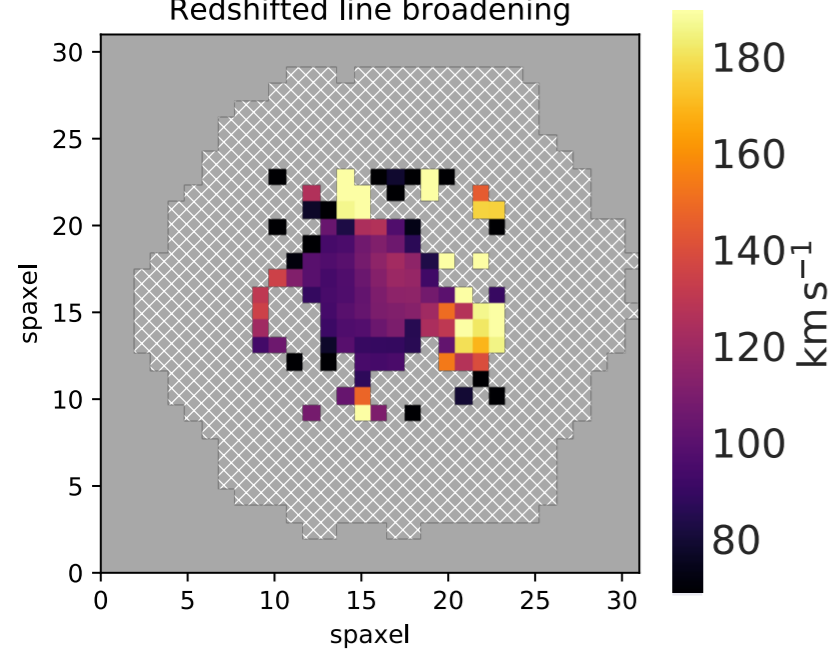
Blueshifted velocity



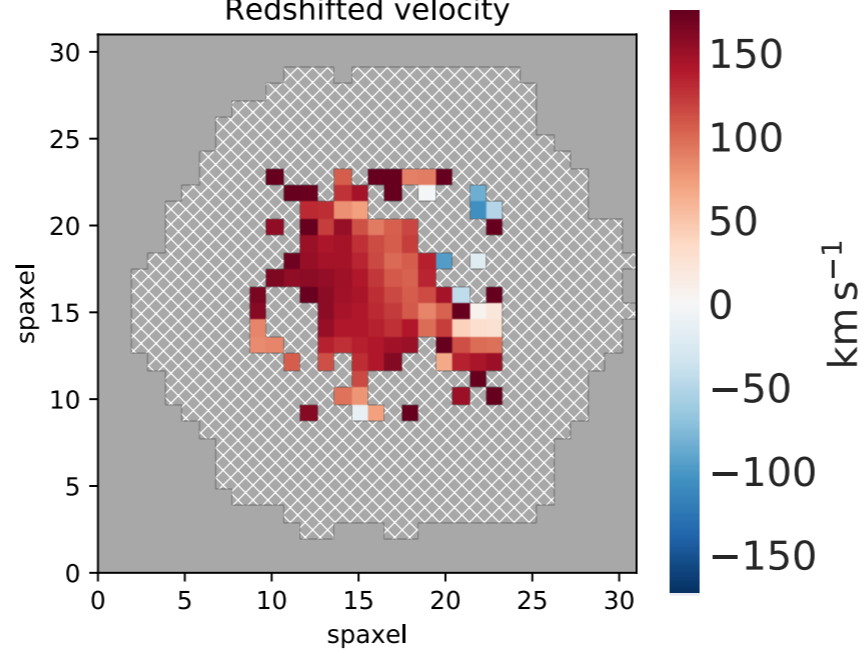
Blueshifted line flux



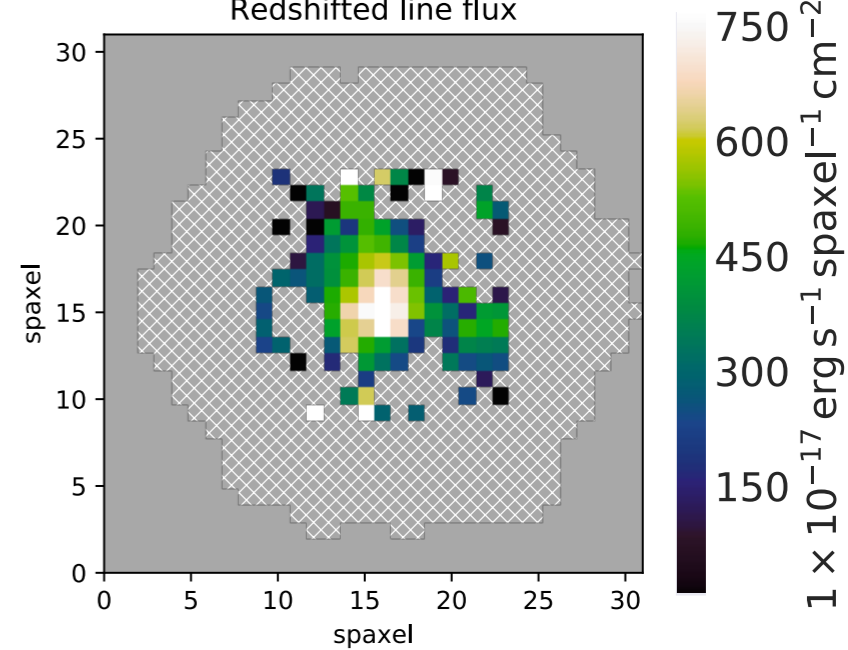
Redshifted line broadening



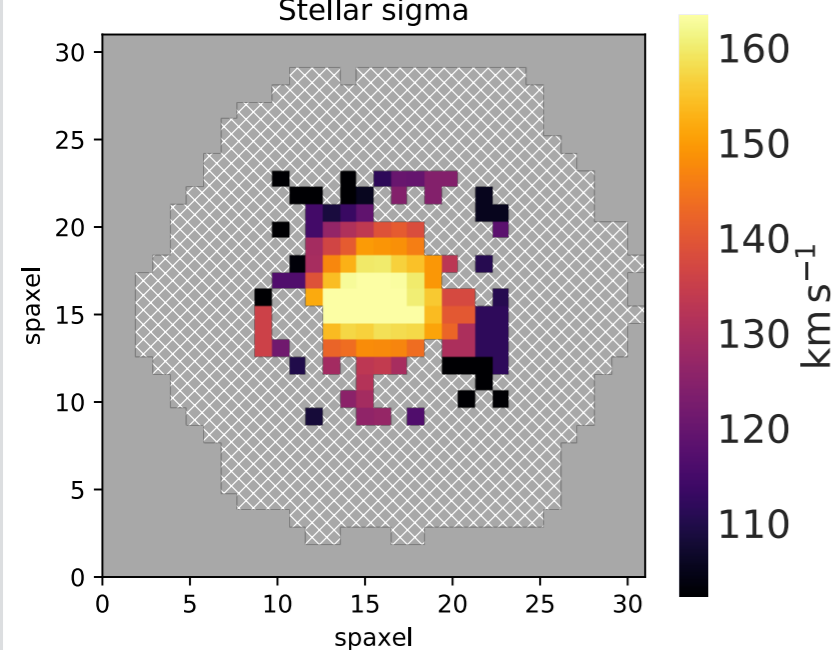
Redshifted velocity



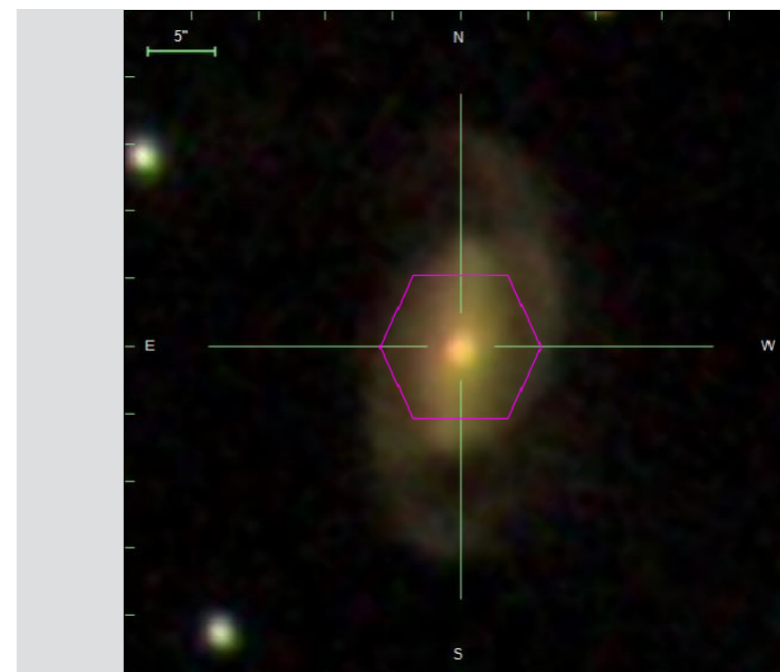
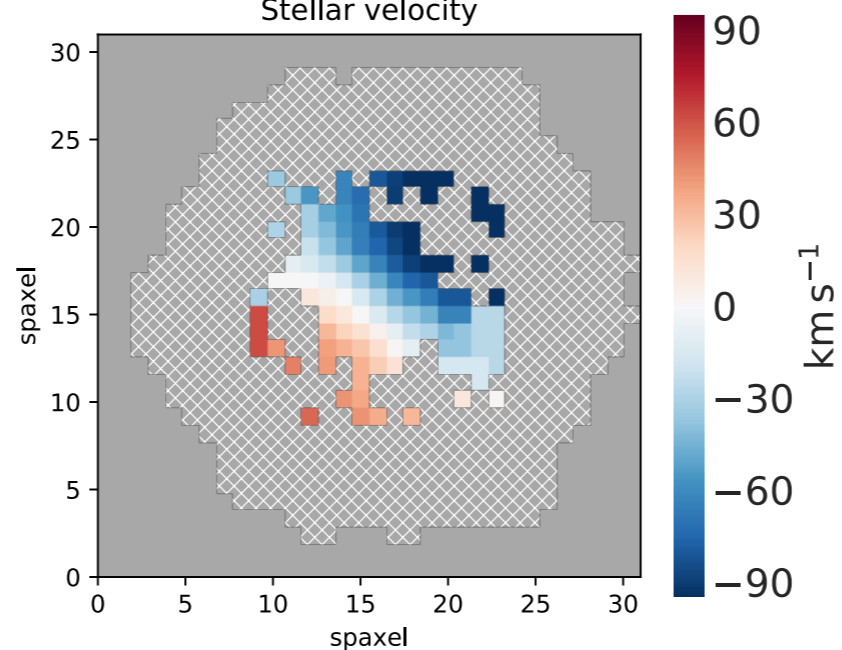
Redshifted line flux



Stellar sigma



Stellar velocity



Summary

Double-peaked galaxies: 0.6% of the SDSS DR7 galaxies

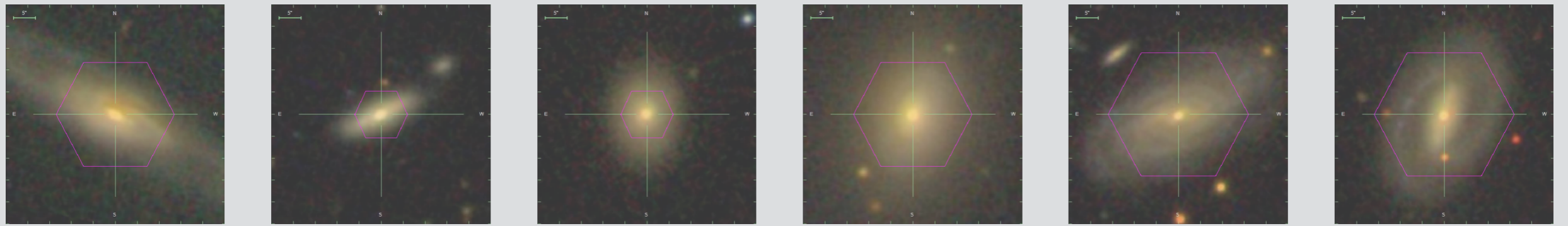
Surprising high fraction of S0 : mergers?

MaNGA data mining: using the DAP, 2D maps show standard velocity fields

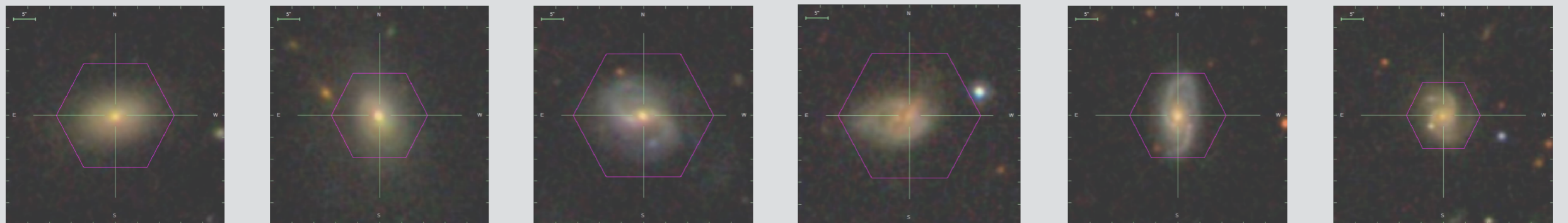
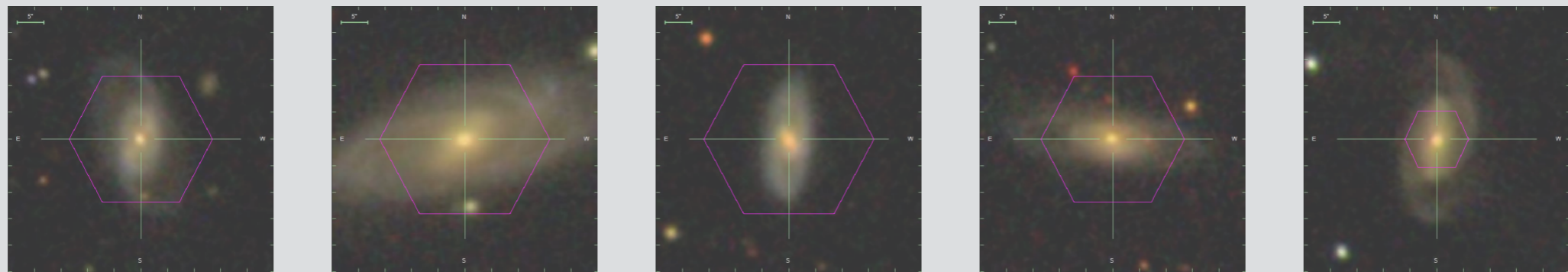
Double-peaked features fitting: one μ , one σ for all lines

How does the emission line maps look based on the double-gaussian fits?

What are we going to learn from molecular gas observations?



Thank you for your attention



Sample classification based on SDSS spectra

