

21-cm signal from Reionization : Astrophysical parameters reconstruction using Supervised Learning

Action Fédératrice

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Context

Models :

- Theoretical
- Numerical
- ...

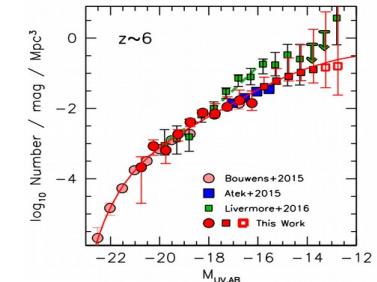
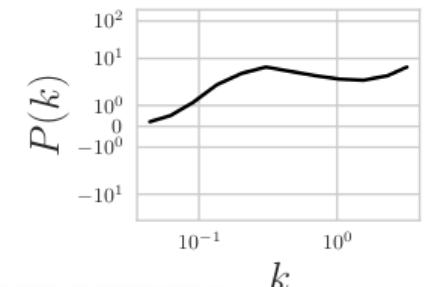
Parameters :

ζ_{ion}
 M_{\min}
 R_{mfp}
 T_{vir}

Reconstruction:

- Bayesian MCMC
- Supervised Learning
- ...

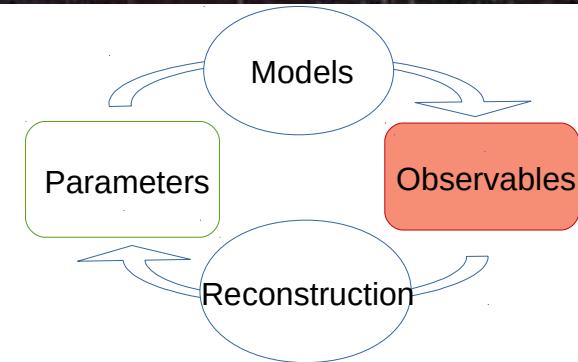
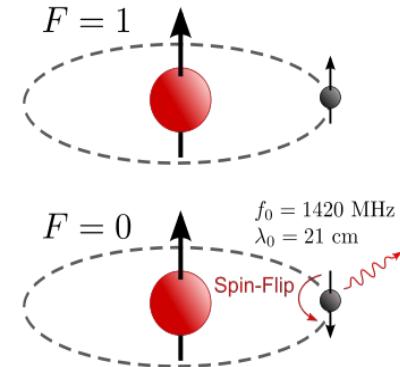
Observables :



Bouwens et al.
2017

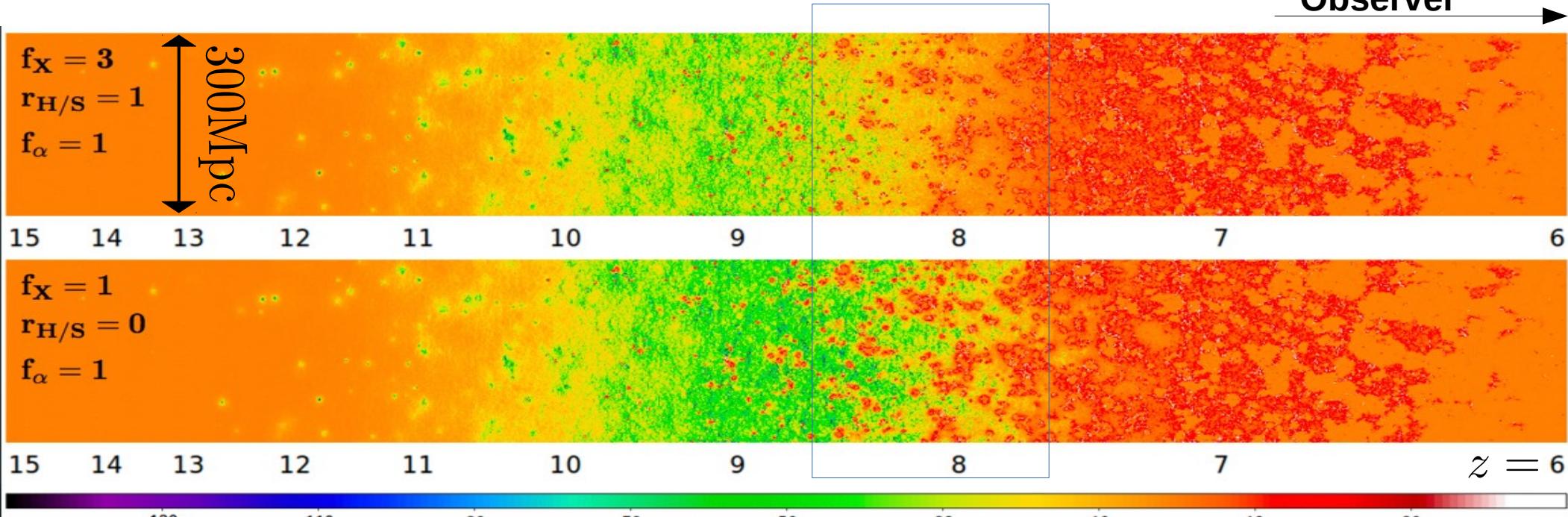
21cm Signal

- Determined by complex correlated physical effects



$$\delta T_B \propto 28(1 + \delta)x_{\text{HI}} \left(\frac{T_S - T_{CMB}}{T_S} \right) \left(1 + \frac{1}{H} \frac{d\nu}{dr} \right)^{-1} \text{mK}$$

Observer →



Semelin et al. 2017

Semi-numerical code 21cmFAST

Mesinger et al. 2011

21cmFast

Parameters :

Ionizing emissivity

ζ_{ion}

Lyman-Limit system
influence

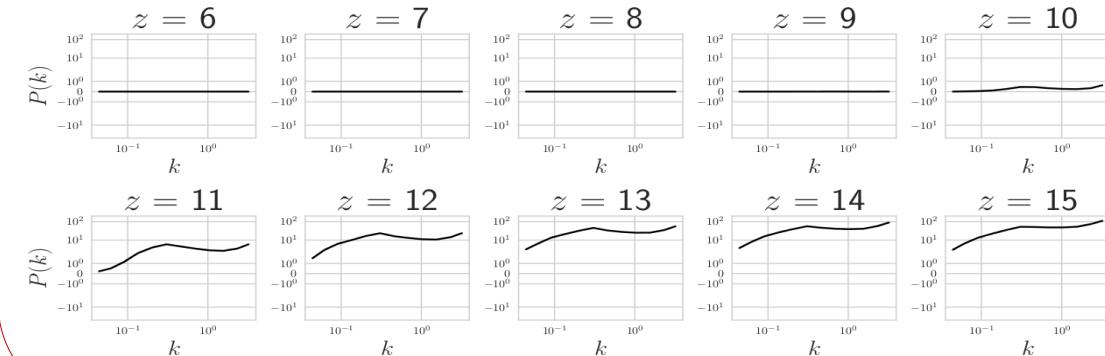
R_{mfp}

Minimal mass of proto-
galaxies

T_{vir}

Observables :

21cm power spectrum



Neural Networks

Parameters reconstruction using supervised learning



Analysing the 21 cm signal from the epoch of reionization with artificial neural networks

Hayato Shimabukuro^{*} and Benoit Semelin^{*}

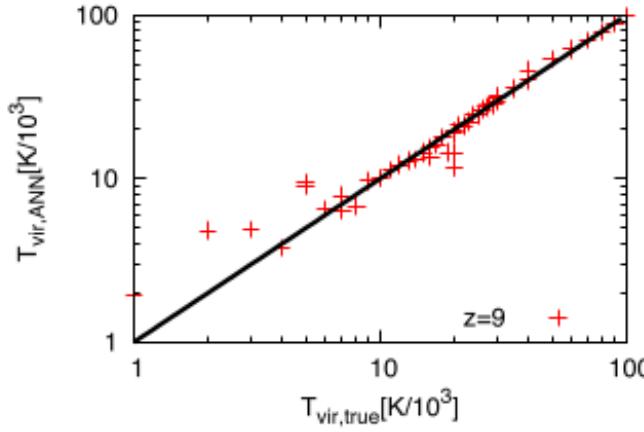
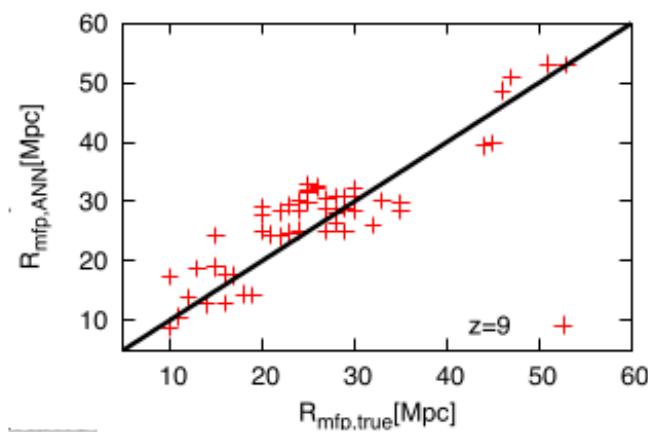
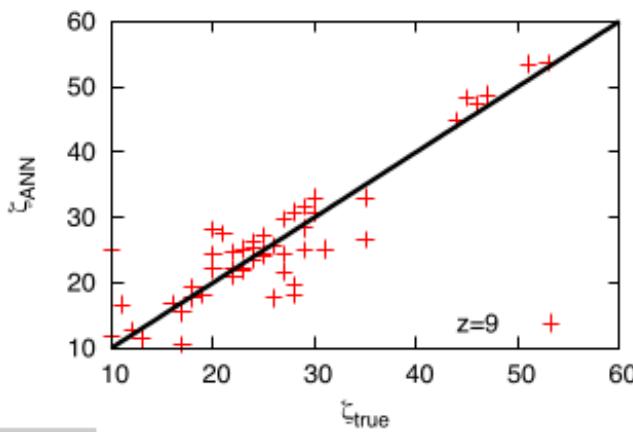
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ABSTRACT

The 21 cm signal from the epoch of reionization will be one of the first observational signatures of the Epoch of Reionization. While a simple model can be used to predict the signal, it is not able to account for all physical processes. In this paper, we show that a neural network can be trained to predict the signal from the underlying astrophysical parameters. We find that the neural network can predict the signal with a high accuracy, even when the input parameters are not well constrained. This method can be used to extract information about the underlying astrophysical processes from the observed data.

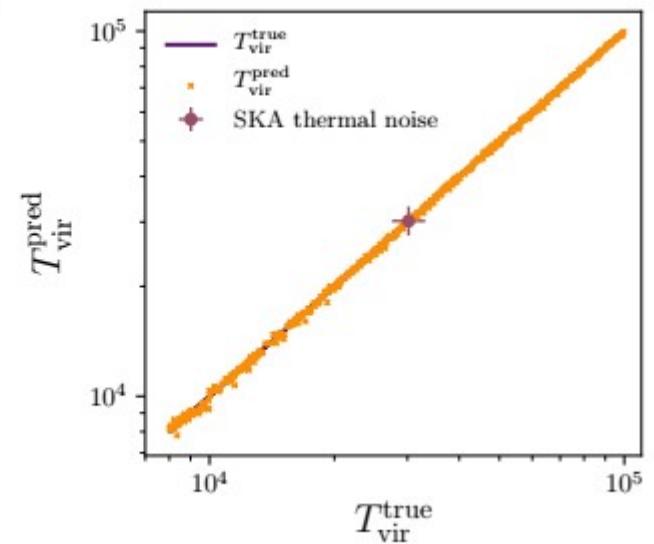
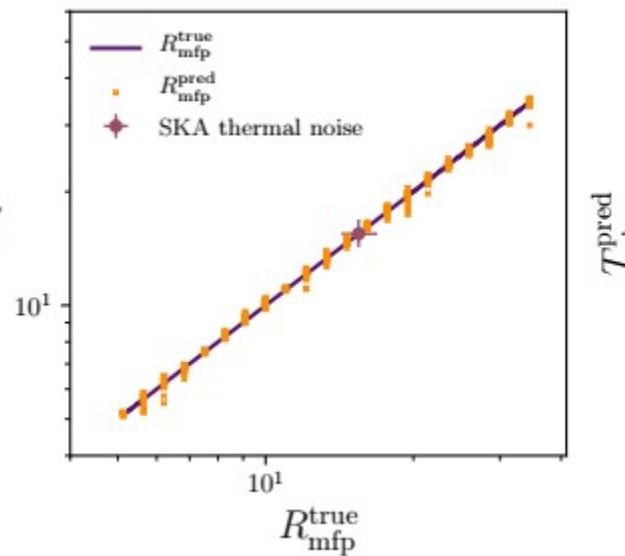
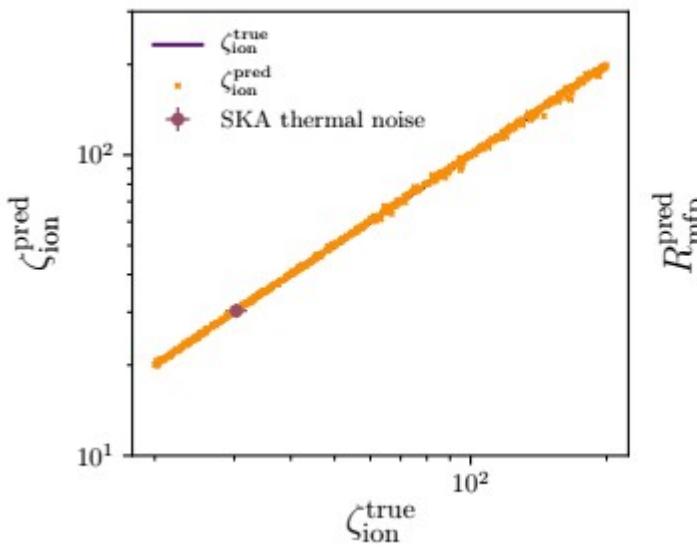
	$\chi_{\text{wo/noise}}$	$\chi_{\text{w/noise}}$	$\chi_{\text{w/noise, zevolution}}$	$\chi_{\text{w/noise, reduced}}$
R_{mfp}	0.228	0.258	0.172	0.262
ζ	0.271	0.288	0.168	0.290
$\log(T_{\text{vir}})$	0.027	0.038	0.019	0.029



the quality of the reconstruction and that using the power spectrum at several redshifts as an

Improved Neural network

- 14 vs 80 neurons, 70 vs 2400 points in the learning samples, redshift evolution
- Better determination of the learning parameters



$\chi^2_{\text{wo/noise}}$	
R_{mfp}	0.228
ζ	0.271
$\log(T_{\text{vir}})$	0.027

$\times 10$ accuracy

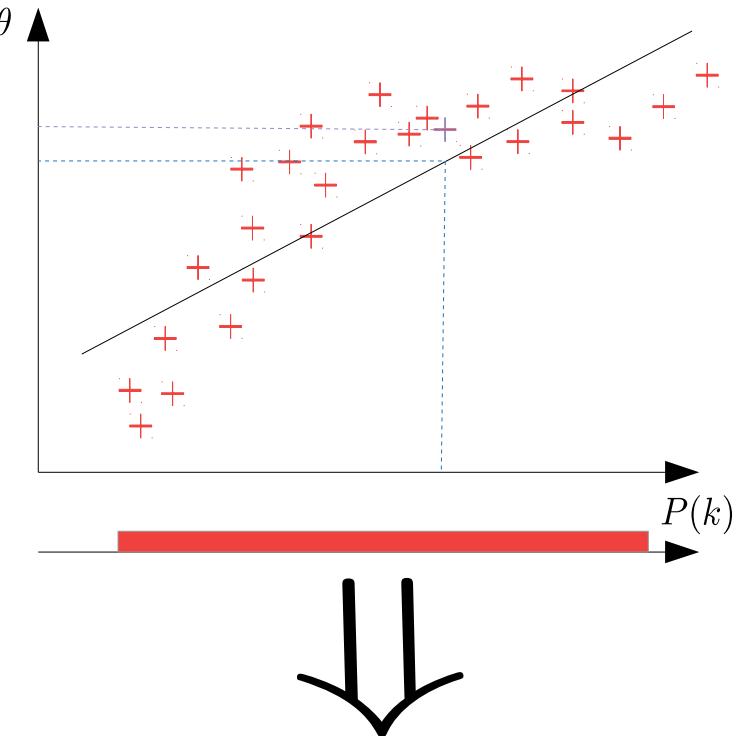
	χ
R_{mfp}	0.02526
ζ_{ion}	0.01369
$\log(T_{\text{vir}})$	0.00154

Linear Regression

- Another supervised learning method
- Minimization of :

$$\min_{\alpha, \beta_j} \sum_{i=1}^{N_{\text{Sample}}} \left(y_i - (\alpha + \sum_{j=1}^{N_{\text{dim}}} \beta_j x_{i,j}) \right)^2$$

$T_{vir}, \zeta_{ion}, \dots$

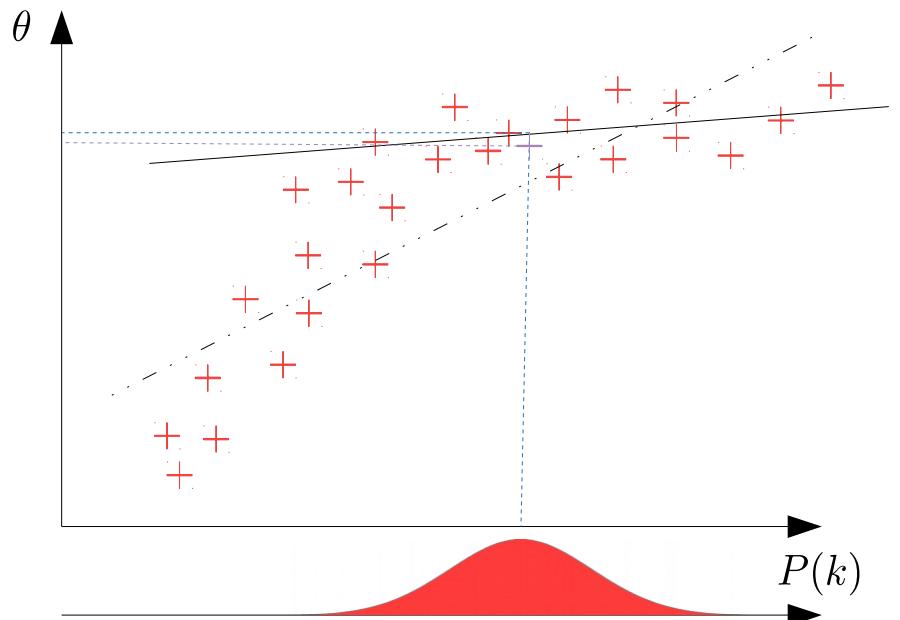


120 dimensions !

	ζ_{ion}	R_{mfp}	$\log(T_{vir})$
$\chi_{\text{Shimabukuro}}$	0.271	0.228	0.027
$\chi_{\text{Neural Network}}$	0.01369	0.02526	0.00154
$\chi_{\text{Least Square}}$	0.01816	0.07996	0.00294

Kernel Regression

$$\min_{\alpha, \beta_j} \sum_{i=1}^{N_{\text{Sample}}} K_\sigma(x_0, x_i) \left(y_i - (\alpha + \sum_{j=1}^{N_{\text{dim}}} \beta_j x_{i,j}) \right)^2$$



	ζ_{ion}	R_{mfp}	$\log(T_{\text{vir}})$
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$\chi_{\text{Kernel Regression}}$	0.01192	0.06129	0.00275

Ridge Regression with Kernel

$$\min_{\alpha, \beta_j} \left[\sum_{i=1}^{N_{\text{Sample}}} K_\sigma(x_0, x_i) \left(y_i - (\alpha + \sum_{j=1}^{N_{\text{dim}}} \beta_j x_{i,j}) \right)^2 + \lambda \sum_{j=1}^{N_{\text{dim}}} \beta_j^2 \right]$$

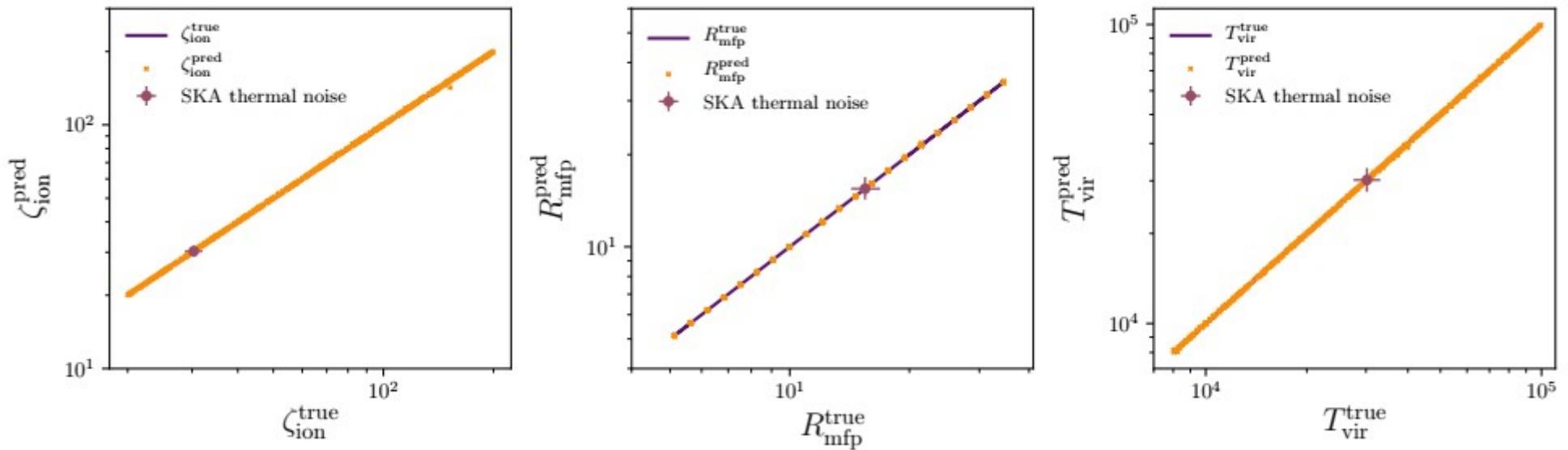
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$\chi_{\text{Kernel+Ridge Regression}}$	0.00676	0.01756	0.00091

Local determination of the hyperparameters

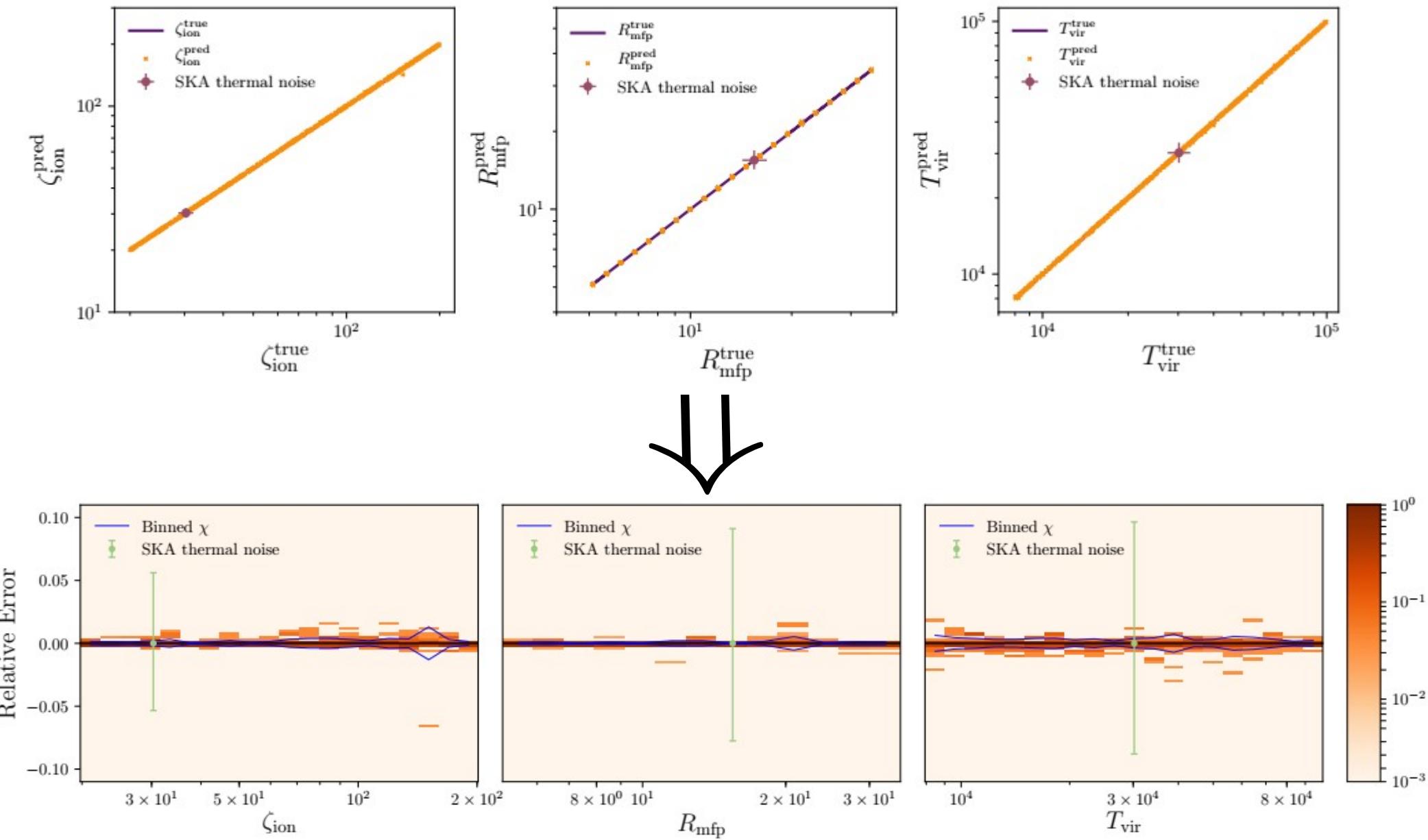
$$\min_{\alpha, \beta_j, \sigma, \lambda} \left[\sum_{i=1}^{N_{\text{Sample}}} K_\sigma(x_0, x_i) \left(y_i - (\alpha + \sum_{j=1}^{N_{\text{dim}}} \beta_j x_{i,j}) \right)^2 + \lambda \sum_{j=1}^{N_{\text{dim}}} \beta_j^2 \right]$$

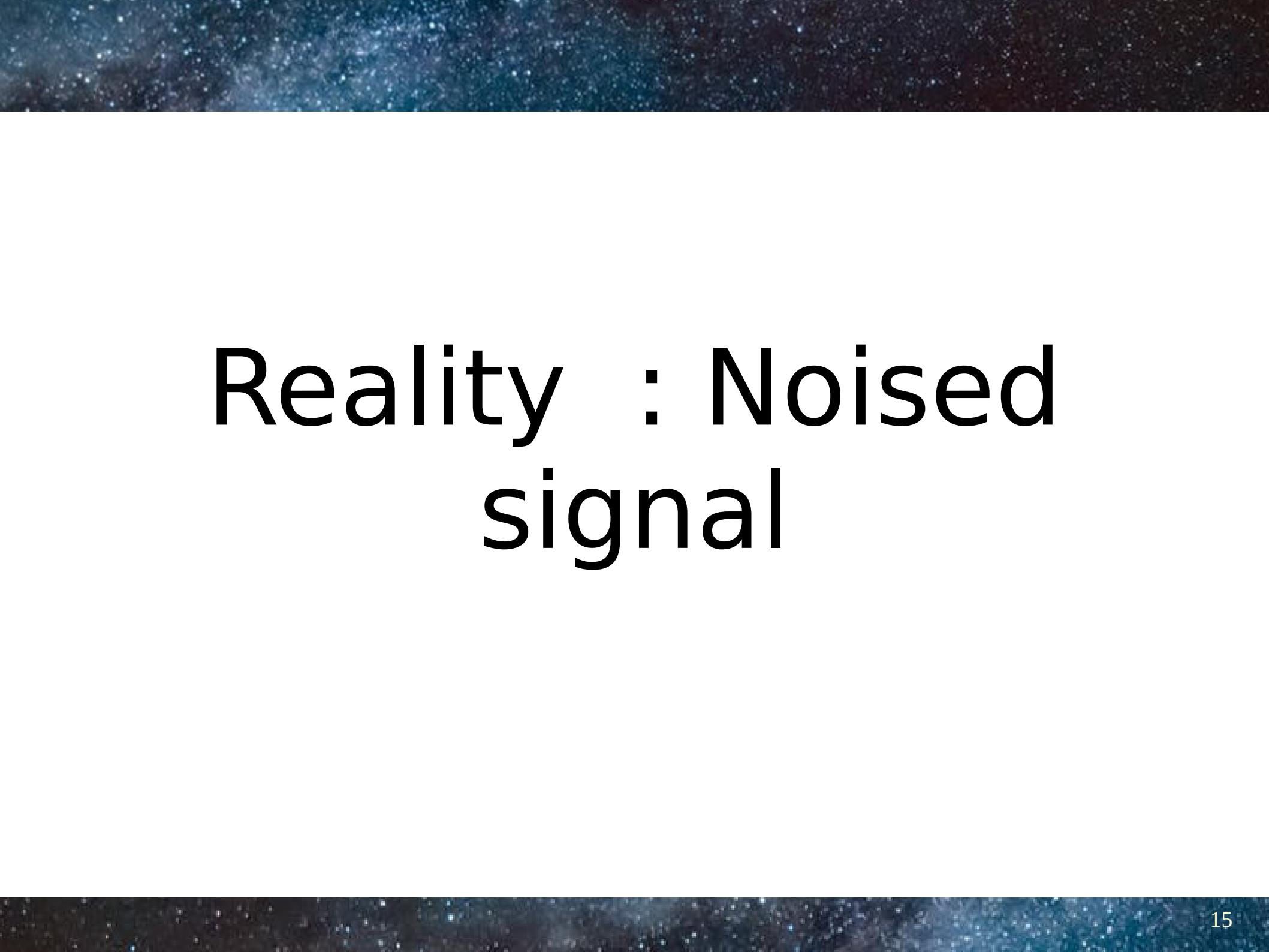
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$\chi_{\text{Kernel+Ridge Regression}}$	0.00676	0.01756	0.00091
$\chi_{\text{Local Kernel+Ridge Regression}}$	0.00463	0.00722	0.00048

Local determination of the hyperparameters



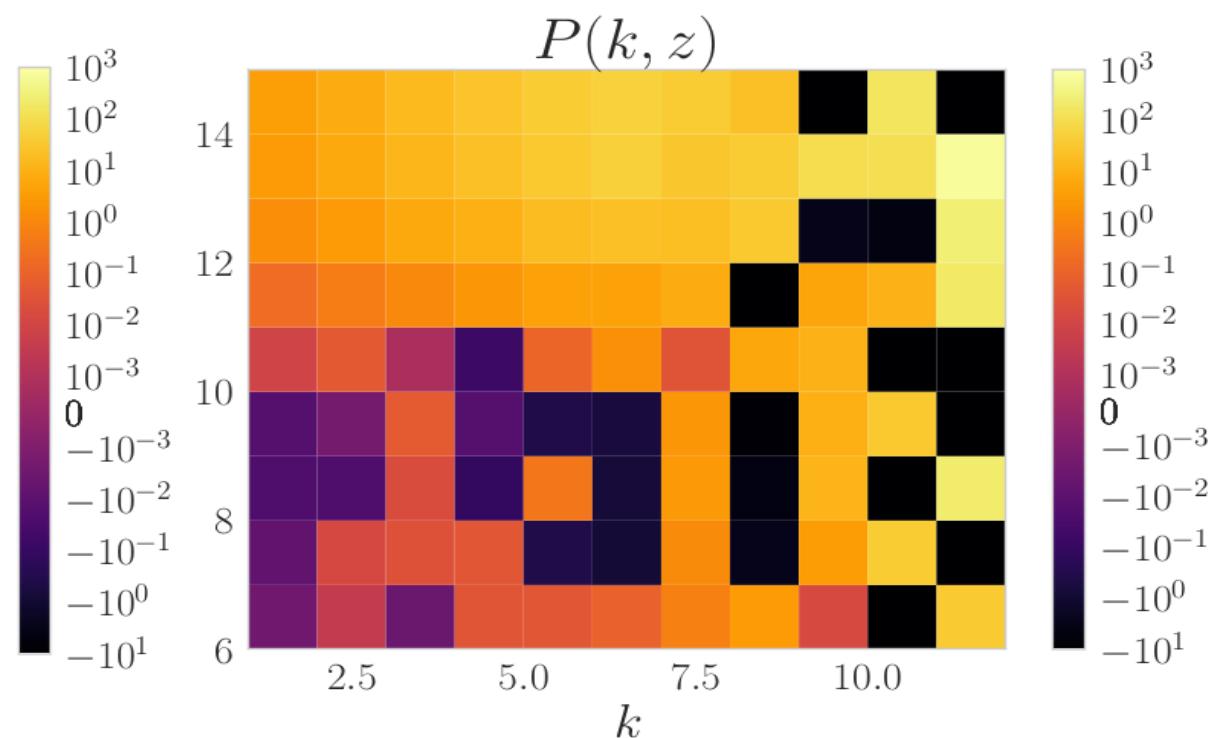
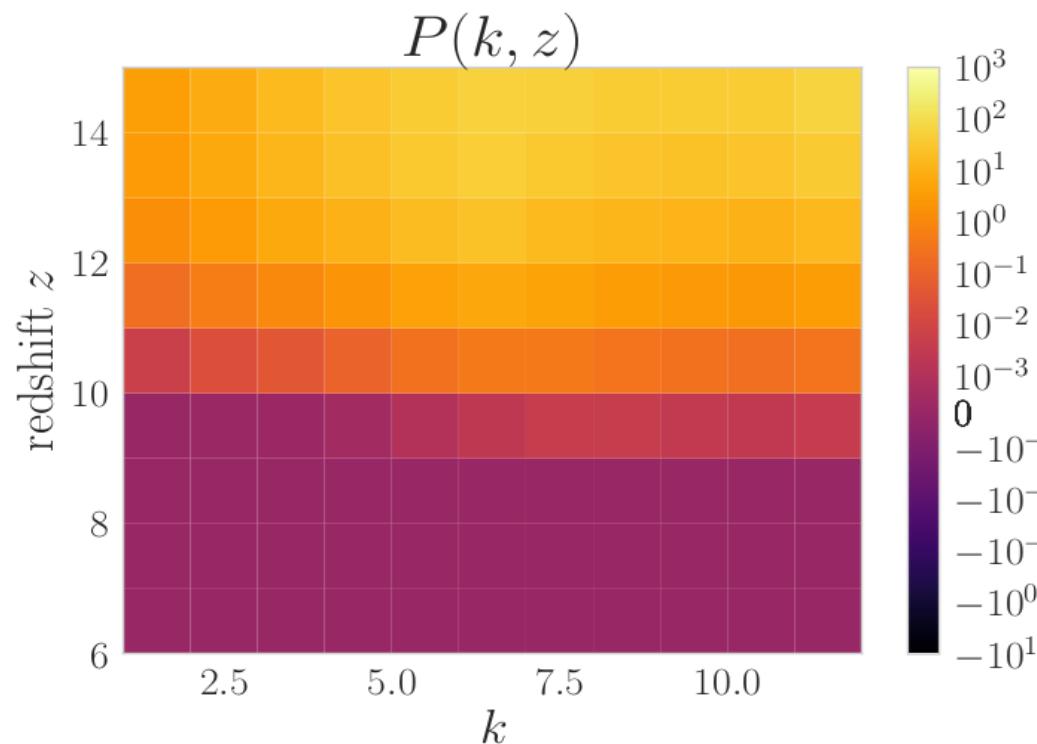
Local determination of the hyperparameters





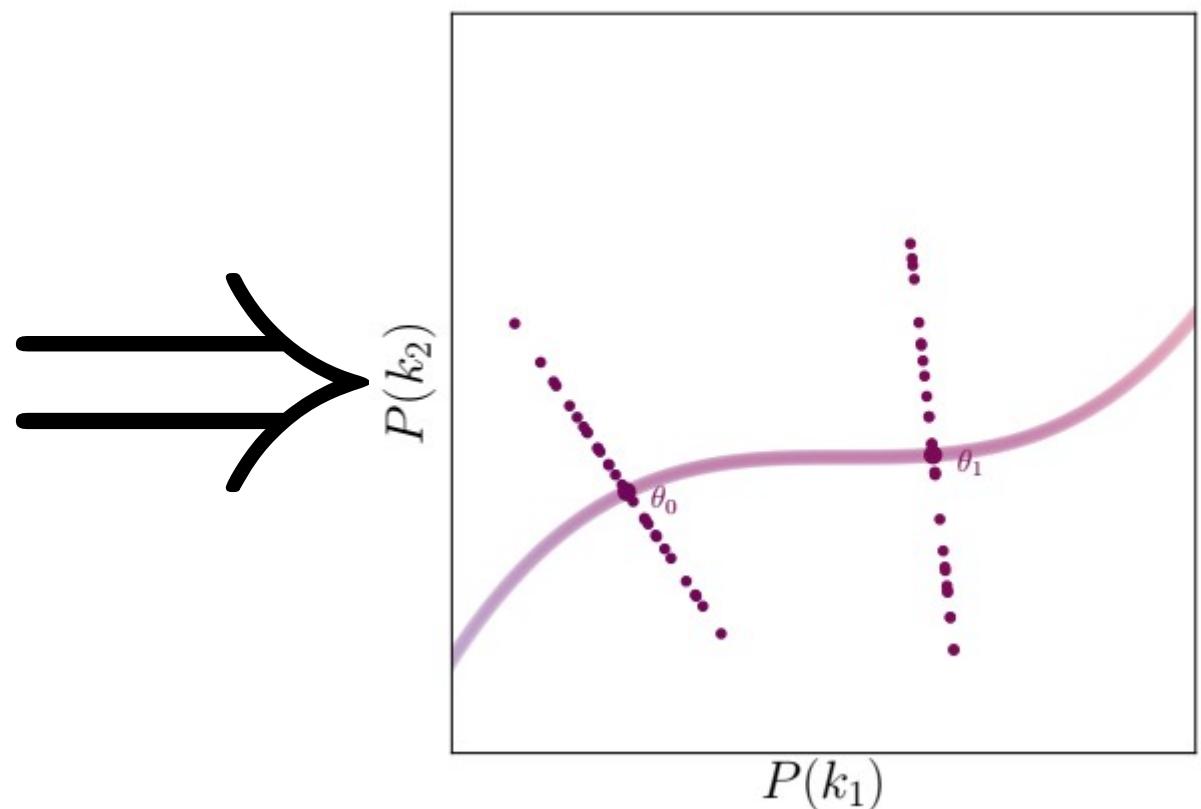
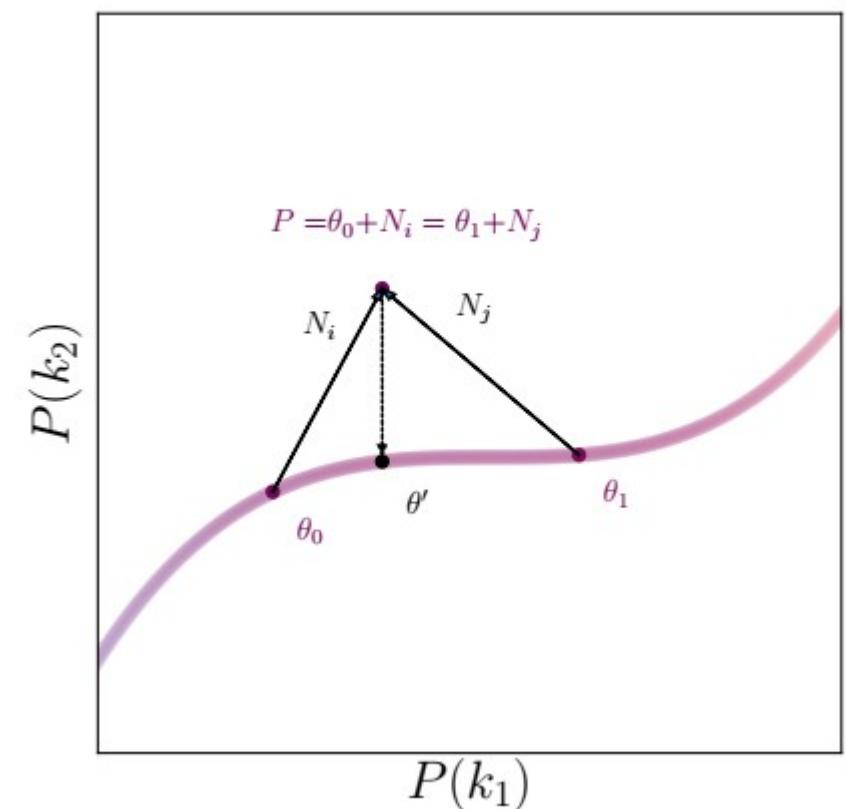
Reality : Noised
signal

Effect of the noise on the input



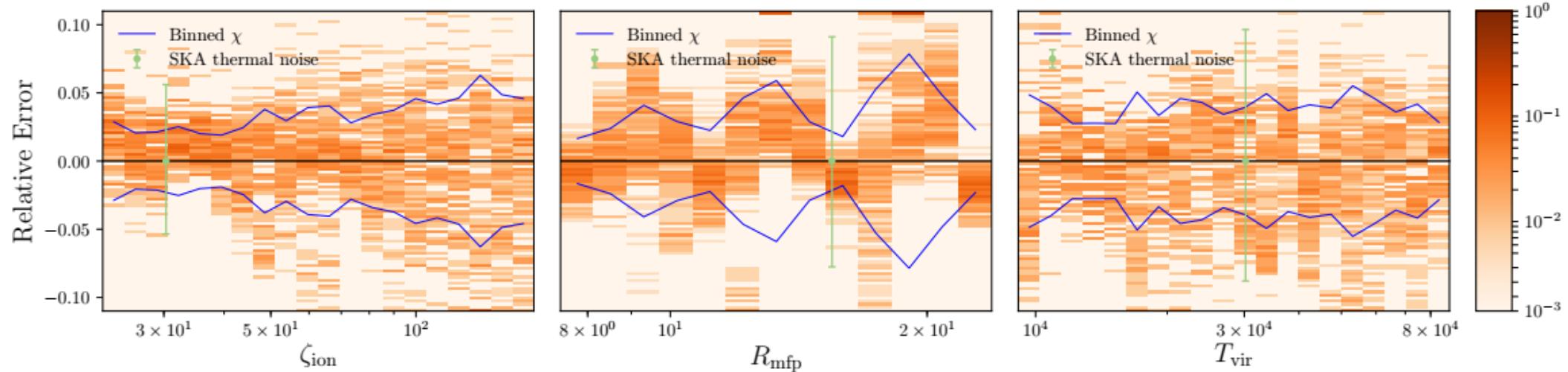
Data Preparation

- Use Signal to Noise Ratio
- Choice of correct answer (maximum likelihood)



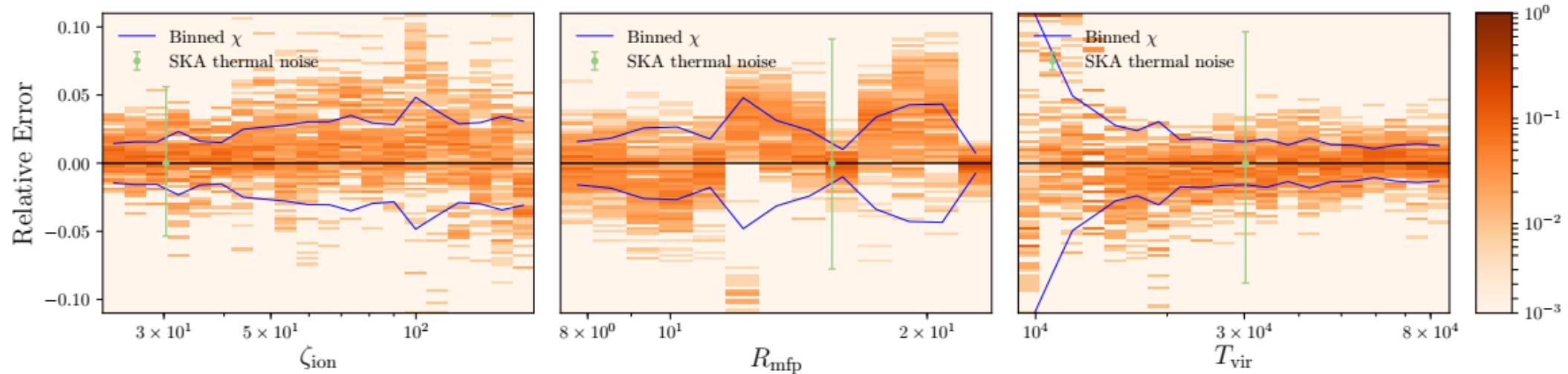
Best Results to this day

With noise	$\chi_{\zeta_{\text{ion}}}$	$\chi_{R_{\text{mfp}}}$	$\chi_{\log(T_{\text{vir}})}$
Shimabukuro 2017	16.8×10^{-2}	17.2×10^{-2}	1.9×10^{-2}
Neural Network	3.70×10^{-2}	4.04×10^{-2}	0.41×10^{-2}



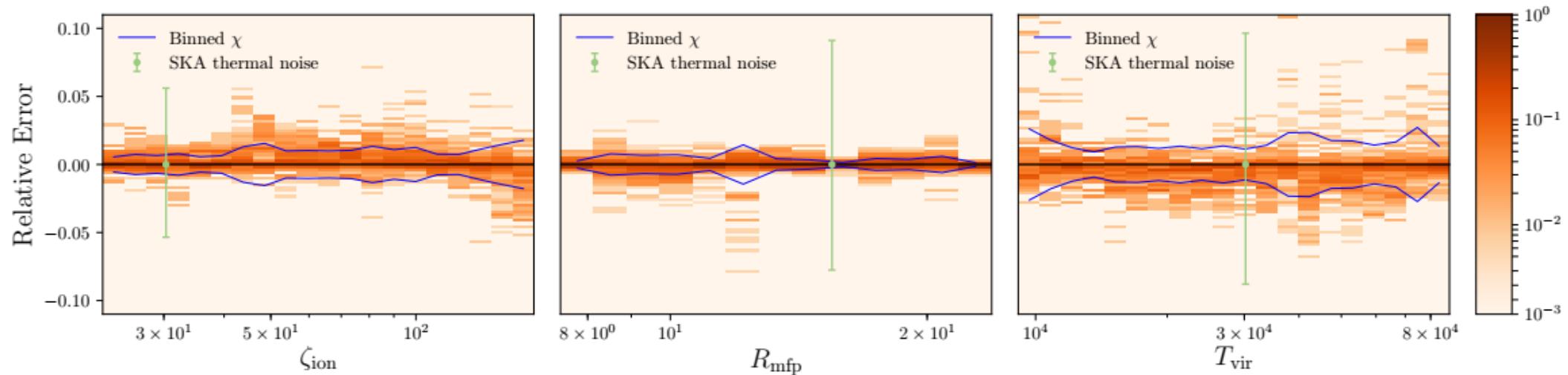
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Global Ridge Kernel Regression	2.88×10^{-2}	2.84×10^{-2}	0.34×10^{-2}



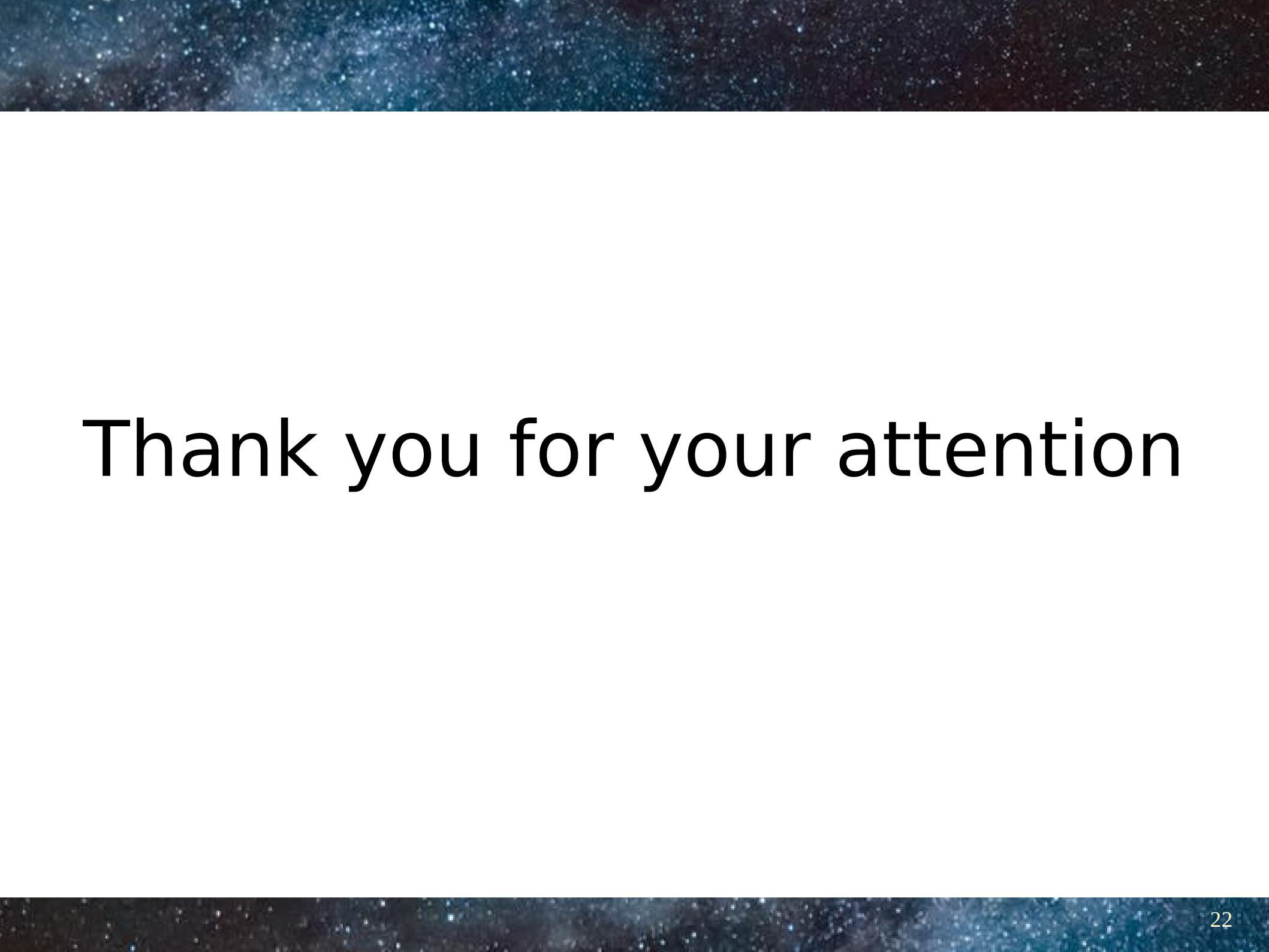
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Global Ridge Kernel Regression	2.88×10^{-2}	2.84×10^{-2}	0.34×10^{-2}
Local Ridge Kernel Regression	1.10×10^{-2}	0.60×10^{-2}	0.16×10^{-2}



Conclusion

- For Noise-free signals :
 - Improvement of the quality of the predictions by a factor **50**
 - Reconstruction error <1 % (SKA bayesian 1-sigma : ~5 %)
- For Noised signals :
 - Improvement of the quality of the predictions by a factor **10**
 - Error on MaxLikelihood value ~1 % (SKA bayesian 1-sigma : ~10 %)
- What's next :
 - Apply to numerical simulation results
 - Bayesian Neural Networks



Thank you for your attention