### ARTS: overview with focus on ISMAR

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# Short history

#### Version 1

- started 2000
- "traditional" unpolarised 1D model
- developed until  $\approx$  2005
- still used for Odin-SMR operational processing

### Version 1.1/2.0

- started 2002
- version 2.0 released 2011
- two main motivations
  - tomographic limb sounding inversions
  - microwave ice cloud observations



Full polarisation (1-4 Stokes elements)

- 1D, 2D or 3D atmosphere
- All observation geometries covered
- Scattering of thermal emission
- Support practical retrievals
- Focus on microwave measurements
  - basic support for radio occultation
  - single scattering radar simulator
  - IR, at least clear-sky works fine

### Key aspects

#### Flexible

important for users

### Modular

important for developers

#### Documented

important for all

#### Maintained

important for users

### Open source

important for all





### 3 Scattering

Define/calculate gas absorption and surface properties

- Pencil beam monochromatic radiative transfer
  - ray tracing
  - integrate along propagation paths
- Inclusion of sensor characteristics
- Adopt to calibration scheme
  - set iy\_unit to "RJBT" before calling yCalc

## Calculation of absorption

- Line-by-line calculation based on
  - ▶ formats: HITRAN, JPL, LBLRTM and ARTS
  - line mixing can be considered (O2 @ 60 GHz)
- Absorption models and continua terms
  - gases: Liebe/Rosenkranz, CKD, HITRAN-CIA ...
  - Iiquid water: MPM93 and Ellison07
- On-the-fly or using lookup table
- All you need for ISMAR!?

- Other features
  - Zeeman and Faraday rotation
  - other planets handled

# Surface radiative properties

Upwelling radiation calculated as

$$I_u = I_e + \sum_i R(\theta^i) I_d(\theta^i)$$

- Workspace methods for flat surface
  - constant or frequency varying scalar reflectivity, following given refractive index
- Lambertian surface
  - so far just a basic method
- FASTEM
  - don't use for incidence angles > 80° or above 400 GHz
- External data
  - possible to import and interpolate fields of R
- Output of EUMETSAT study should be implemented
  - What else could be needed for ISMAR?

## The surface and 3D

- Possible to operate with surface types (land-sea mask)
- An example (with resolution actually used):



A rough surface topography can be specified
Resolution follows the atmospheric lat and lon grids

Complete measurement vector calculated as

 $\mathbf{y} = [\mathbf{Hi}_1; \mathbf{Hi}_2; \dots; \mathbf{Hi}_n]$ 

#### The response matrix H can include effects of

- polarisation response, antenna pattern, sideband weighting, backend channel response, frequency interpolation . . .
- ► The approach assumes that **H** is identical for each view,
- but not valid for AMSU-A, ISMAR ....

as polarisation response changes with viewing angle

# "Met MM system"

- Calculates H for a scan sequence
- DEIMOS, MARSS and ISMAR included
- Antenna response not yet handled
- Perfect balance between sideband assumed
- Rectangular channel response assumed
- Number of frequencies per passband can be set
- Scan positions to include can be selected
- Setting of sensor\_los
  - zenith angle: direction of instruments reference direction (177° if no pitch or roll)
  - azimuth angle: direction of scan plane (flight direction of no roll)
- Use/mimic Atmlab's ismar\_los.m

#### Test of polarisation part Flight B893, processing and figures by Stuart Fox



### Simulation of complete flights Flight B893, data close to nadir



- ERA-Interim (0.7°) with LWC considered (purely absorbing)
- 3D + Met-MM + surface set as:
  - ocean: FASTEM
  - Iakes: FASTEM, with reduced wind speed (and zero salinity)
  - land: r = 0.15

# OEM type retrievals

- Jacobian (weighting functions) provided for a number of atmospheric and sensor variables
- but not yet anything useful for surface reflectivity
  - something basic will be added soon

- Running OEM so far requires Atmlab
- but OEM inside ARTS is work in progress







## Single scattering properties

- Must be provided in a ARTS specific format
- Horizontally aligned particles handled
- ARTS comes with Mishchenko's T-matrix code
  - probably less useful than expected (see Eriksson et al. AMT, 2015)
- Liu and Hong DDA databases are at hand in ARTS format
  - critical limitations of both databases
  - just total random orientation considered
- A new DDA database has been started
  - more in presentation by Robin

# Setting of particle number densities

The core part is blind to particle type

- For each particle type you need to specify
  - single scattering
  - particle number density (PND) field [#/m3]
- The PND field can be set directly
- and now also based on hydrometeor profiles, combined with user-specified PSD:



# Solution methods

- Discrete Ordinate Iterative method (DOIT)
  - determines the complete radiation field in an iterative manner
  - 1D and 3D versions
  - required settings need some care
  - can provide weighting functions (very slow option) >
- Monte Carlo (MC)
  - follows "photons" backwards in a random manner
  - only 3D
  - easy to use

Both methods are fully polarised (vector RT)

Present recommendation for ISMAR: DOIT-1D



### Discussion

- Neither DOIT or MC is very fast! But sufficiently fast?
- If no, is there any option?
- Vector radiative transfer is required!
  - V and H radiation streams can not be treated separately when multiple scattering is significant
- Do we need 3D scattering calculations?
  - for ICI probably yes
  - less clear for ISMAR
- A master project on the subject has been started

### Conclusions

ARTS has the basic capabilities to handle ISMAR

- Main lack of data:
  - surface properties above 200 GHz
  - any single scattering data that are OK from 90 to 664 GHz?
  - single scattering data for horizontal alignment
- Do we need a faster scattering method?
  - should then be polarised
- Do we need to consider horizontal structures?
  - that is, scattering in 3D