# ARTEMIX

### ALMA REMOTE MINING EXPERIMENT



EUROPEAN ARC ALMA Regional Centre || IRAM



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Archive and Data Mining



ARTEMIX

### Goals

(i) Search by products not by instrumental configuration(ii) Provide trans-project queries (ie famous sources)(iii) Have a rapid idea of the data content (fits files)

### Means

(i) ALMA observing configuration previews (meta-data)(ii) ALMA cube previews (science products QA2)

## Context (1)

(i) To stand just beyond what is provided by the ALMA Observatory Science Archive (not delivering data, not providing material for data reduction). —> Redirection to the ALMA Science Archive

(ii) To use public meta-data and public fits data cubes

(iii) Not to redo what already exist in the ASA (ie rapid metadata query by multi-filters)

## Context (2)

### Other similar software tools

- A powerful server-side visualization tool (PI: Erik Rosolowsky) which allows users to browse and manipulate the very large ALMA data cubes without having to download them to disk first : CARTA (Cube Analysis and Rendering Tool for Astronomy)
- Japanese Virtual Observatory (JVO) science-ready ALMA images (JVO portal (<u>http://jvo.nao.ac.jp/index-e.html</u>)
- ESO dev project (Mansardi) study to determine if re-imaging and delivering the full data cubes is feasible (data reduction / QA / storage) + eventually plug automated analysis tools (like ADMIT) —> long term plan

## Context (3)

—> To Provide **a pilot study** of **remotely** operated tools for **quick look** visualisation (regular discussions with F. Stoer to keep informed about the current developments

—> Not developed in ESO coding standards, but with selected techno i.e. Serveur HTTP:NodeJS, Database:MongoDB; FITS server:python...

—> Developed at the Paris Observatory, LERMA, in the framework of the French AA-ANO3 duties coordinated by the OASU (OASU, Obs. Paris, OSUG, IRAM)

## 1- Sources Display

—> Very light information. Ingested in a Mongo Data Base on a daily basis

—> Display the sky coverage (RA, DEC) via the plotly library (interactive : zoom, name and position)

—> Basic search method (by sesame-name, alma-name or position inside a given radius) : resolve the source position when necessary, grab pointings (projects) around the object if the source is nearby and extended

**Goal** : retrieve all the observed projects for a given source (region) of interest (whatever the PI, whatever the project)

#### 08 Q+00 DEX# 1 = 🔜 👗



Distribution of ALMA sources (2010-2016)

Right Ascention [hours]

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## 2- Region of Interest

—> Display the **frequency range observed** (basebands) for all the projects that correspond to a given region of interest

--> Display the region of interest into **AladinLite** (help getting velocity / redshift if searched by sesame-name)

—> Display a **table of the header keywords** for all fits files that correspond to this region of interest (different Project codes, different FoV...)

—> (Soon) : display a table with all the metadata for these projects (resolution, t\_obs..)

**Goal** : quick visual inspection of what has been observed .vs. what has been imaged. Provide a link to the data cube (or 2D if in yellow)

—> On click in the table : overlay the frequency slice that has been imaged (**metadata vs fits header**), overlay the FoV (fits header box)

—> Link to **ADS** for publication check

--> Link to ESO/NRAO ALMA archive for data retrieval

#### Home page Tools - About Help Admin

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089+0902×1===.





| Show all data. Initially shown are the data of which the f | lename ends with ' | ".pbcor.fits", | ".pbcorr.fits", | ".image.fits", | "line.fits", | "cont.fits" o | r "clean.fits" |
|--|--------------------|----------------|-----------------|----------------|--------------|---------------|----------------|

|   | Fits file | Target  | RA          | DEC          | Cube size     | Freq. range     | Proj. code     | uid                  |
|---|-----------|---------|-------------|--------------|---------------|-----------------|----------------|----------------------|
| 1 | info 📼    | NGC1365 | 03:33:36.38 | -36:08:25.70 | 2916x2560x10  | 246.011 245.864 | 2013.1.01161.S | uid://A001/X12f/X321 |
| 2 | Info 📼    | NGC1365 | 03:33:36.38 | -36:08:25.70 | 2916x2560x30  | 229.066 229.506 | 2013.1.01161.S | uid://A001/X12f/X321 |
| 3 | Info 💌    | NGC1365 | 03:33:36.38 | -36:08:25.70 | 2916x2560x1   | 228.458 246.218 | 2013.1.01161.8 | uid://A001/X12f/X321 |
| 4 | info 💌    | NGC1365 | 03:33:36.38 | -36:08:25.70 | 2916x2560x10  | 230.572 230.718 | 2013.1.01161.S | uld://A001/X12f/X321 |
| 5 | info 💌    | NGC1365 | 03:33:36.38 | -36:08:25.70 | 2916x2560x11  | 243.714 243.553 | 2013.1.01161.S | uld://A001/X12f/X321 |
| 6 | into 💌    | NGC1365 | 03:33:36.38 | -36:08:25.70 | 1344x864x500  | 228.4 230.323   | 2013.1.01161.S | uld://A001/X126/X319 |
| 7 | info 🔹    | NGC1365 | 03:33:36.38 | -36:08:25.70 | 216x216x497   | 228.444 230.355 | 2013.1.01161.S | uid://A001/X120/X31b |
| 8 | info 📼    | NGC1365 | 03:33:36.99 | -36:08:36.33 | 2048x1500x121 | 229.064 229.523 | 2013.1.01161.S | uid://A001/X12f/X317 |
| 9 | info 📼    | NGC1365 | 03:33:36.99 | -36:08:36.33 | 2048x1296x1   | 229.837 246.192 | 2013.1.01161.S | uid://A001/X12f/X317 |



## 3- Quick Look Viewer

—> Display the data cube (2 images, 2 spectra) : 1 channel map, 1 moment map, 1 spectra extracted from a pixel, 1 spectra extracted from a spatial region (square). Interactive and self-consistent

—> Based on **GILDAS Mapping « go view »**. Same functionalities implemented (frequency selection, region selection, integrated flux computation)

---> Link to the detailed fits header

**Goal** : give a quick look preview of the data cube content

- Sometimes (oldest data) the x-axis reference frame varies (LSR, HELIO, BARY...). No computation done apart Freq —> Vel and Vel —> Freq in the same frame.
- Only slices have been extracted (not always the best cut-out or spectral resolution). No re-calibration applied.

--> PI are encouraged to retrieve the ALMA raw data and use the standard pipeline



## Volumetry

Data cubes (ie  $512 \times 512 \times 3000$ ) in  $(x,y,v) \rightarrow 10s$  of MB to 10s of GB (16-20 GB at max). Difficult to handle and visualise cubes larger than that anyway. This would need other methods (ie cut-out service)

Note : The fits file are produced in QA2 —> only slices around the line of interest (less than ~10% of the observed (x,y,v)). Here again, PIs are advised to retrieve the raw data from the ALMA Science Archive.

Total (all fits, as of last week) : ~ 6 TB (updated on a daily basis : not yet achieved)

- 3.5 TB de fichiers > 1 GB
- 2.6 TB de fichiers > 2 GB
- 2.1 TB de fichiers > 4 GB

Solution for fast reading : SSD PCI Express. Very fast load of the data. Fast computation (means are parallelized). Slow part : the network transfert 2 x (NxN) and 2 x Nchans

### Next : a cut-out Service ?

### A cut out service to display / analyse only part of the data fits file.

- 1. Spatially (sub-region in the map) to reduce x-axis and y-axis size
- 2. In Velocity/frequency (range of channels).

To do so, there is a need for

• tools to help the user to choose which part to look at

• tools to split the files in order to scan all the data cube (or all the selected area if still too big).

## Local / Remote

### Local

Large range of analysis tools

### But

Need to download all the fits files to be checked (even if no detection)

Speed limited by local computer performances and/or software optimization (for display) —> often need a local server

### Remote

Optimization on dedicated machines (load fits, calculations)

No need to download fits files on local disk (if many and from different projects)

### But

Delay for download (11 MB/s at most) and loading large file (> 2GB)

Limited analysis (yet)

### Overview



### Thanks !